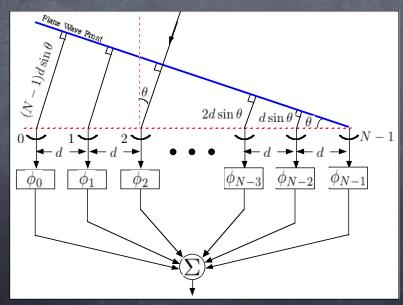


By the time shift theorem of Fourier transforms, we have

$$s(t) \stackrel{\mathcal{F}}{\leftrightarrow} S(f) \implies s(t-\tau) \stackrel{\mathcal{F}}{\leftrightarrow} S(f) e^{-i2\pi f \tau}.$$

So for a sinusoidal or narrowband signal at frequency f_0 , we can replace the delay τ_m by phase shift

$$\phi_m = 2\pi f_0 \tau_m.$$



Assuming narrowband waves and phase shifters with

$$\phi_0 = \phi_1 = \phi_2 = \dots = \phi_{N-1} = 0$$

and N identical elements with effective area $A_e(\theta)$ (gain $G_e(\theta)$) for a wave from direction θ , it can be shown the effective area of the array is

$$A(\theta) = A_e(\theta) \cdot \frac{1}{N} \left| \sum_{n=0}^{N-1} e^{i2\pi n(d/\lambda)\sin\theta} \right|^2$$

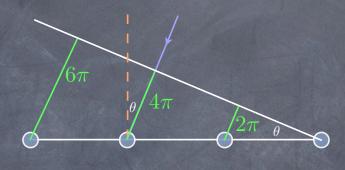
or equivalently

$$G(\theta) = G_e(\theta) \cdot \frac{1}{N} \left| \frac{\sin \left[N\pi(d/\lambda) \sin \theta \right]}{\sin \left[\pi(d/\lambda) \sin \theta \right]} \right|^2$$

Array Length =
$$(N-1)d$$

Larger d implies higer resolution, but there is a price to pay.

 $\underline{\underline{\text{If } d > \lambda/2}}$, we get grating lobes due to constructive interference at Bragg angles:



In order to reduce grating lobes, you must have $d \leq \lambda/2$. You can also

- 1. Use nonuniform spacing of elements;
- 2. Use an $A_e(\theta)$ that reduces the most problematic grating lobes. (elements may be large)

Radio astronomy arrays often have severe grating lobes.

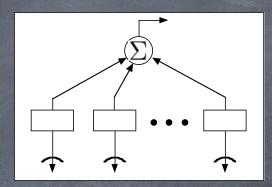
In radar they can be more problematic. Usually take $d \approx 2$

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Usually take $d \approx \lambda/2$.

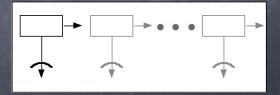
Synthetic Arrays

A real array:



A synthetic array:

Another approach is to use a single element and move it between observations



Signal processing is used to synthesize an "equivalent" array.