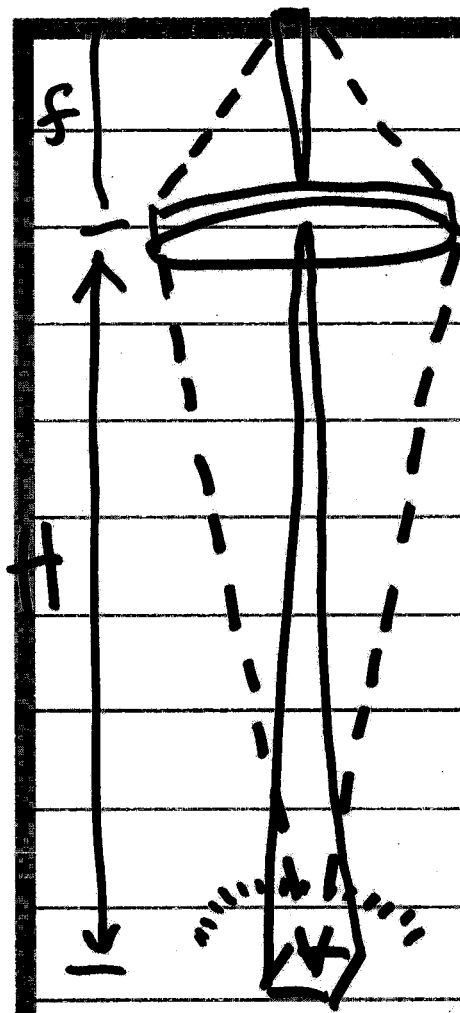


$$I = \frac{\Phi}{\omega} = \frac{\Phi_{\text{total}}}{4\pi \text{ sr}}, \quad 4\pi I = \Phi_{\text{total}}$$

Irradiance @ surface $E = \frac{\Phi}{A}$

$$E = \frac{\cancel{4\pi} I}{\cancel{4\pi} d^2} \leftarrow \text{area of sphere with radius} = d$$

$$E = \frac{I}{d^2}$$



$$A_a = \pi R_a^2$$

$$L = \frac{\Phi}{\omega A} = \frac{\Phi}{\omega s^2} = I$$

$$I = L s^2$$

$$E = \frac{I}{H^2}$$

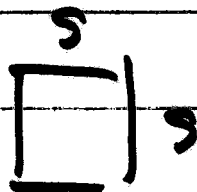
$$E = \frac{L s^2}{H^2}$$

$$E = \frac{\Phi}{A_a}$$

Radiance

$$L = \frac{\Phi}{\omega A}$$

footprint of $(\theta = 0)$
single detector
element



$$\frac{\Phi}{A_a} = L S^2 / H^2$$

$$\Phi = \pi R_a^2 L S^2 / H^2$$

$$\Phi = \frac{Q}{t} \text{ energy/time}$$

$$\frac{Q}{t} = \pi R_a^2 L S^2 / H^2$$

$$Q = t \pi R_a^2 L S^2 / H^2$$

design equation

energy delivered to aperture by a pixel area during time interval t

exp time, integr time

aperture radius

radiance of scene

S : GSD pixel size on ground

altitude

units: watts, meters, seconds, joules, sr

Loss factors

- transit through atmosphere
- quantum efficiency
% of incident photons, which free an electron 20%
- charge transfer

Relate energy Q to photons/electron count

$$Q = \# \text{ photons} \times h\nu$$

$$h\nu = \frac{hc}{\lambda}$$

h : 6.626×10^{-34} Js planck's c.

ν : frequency

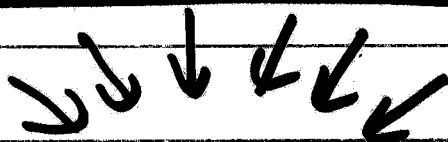
c : speed of light 3×10^8 m/s

$$\lambda : 550 \text{ nm} = 550 \times 10^{-9} \text{ m}$$

(green)

$$Q : \frac{hc}{\lambda} : 3.614 \times 10^{-19} \text{ J / green photon}$$

$$\begin{aligned} M &= \pi L \\ E &= \pi L \end{aligned}$$



dA

assume: radiance
independent of
direction

ISOTROPIC

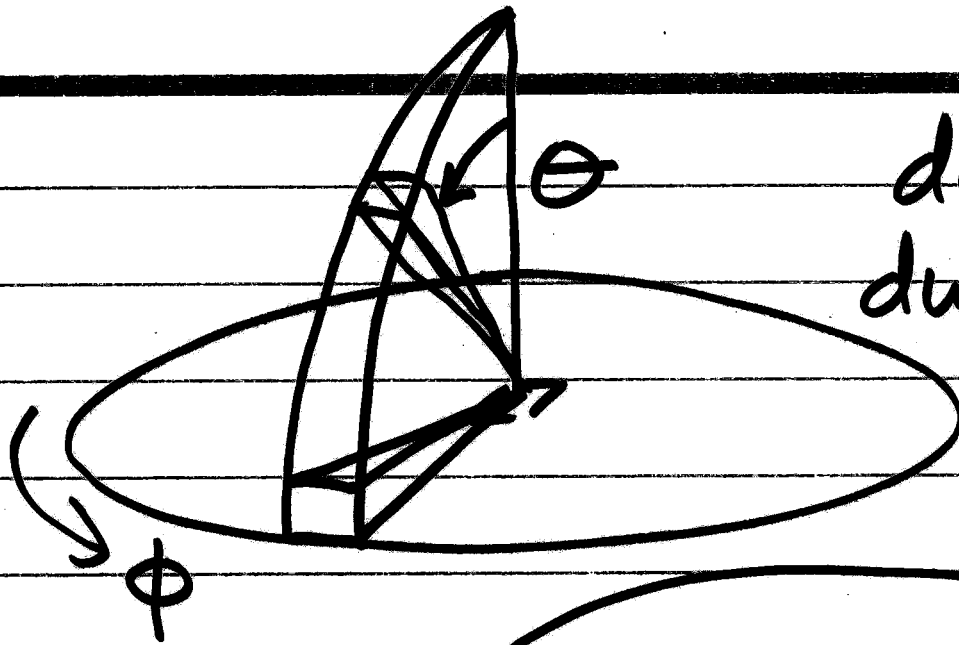
incident power on dA is

(1) proportional to dA (apparent dA)

(2) proportional to $d\omega$ (solid angle)

(3) proportional to strength of EMR: L
radiance

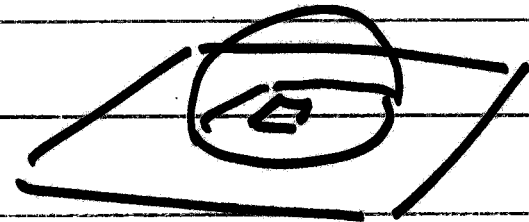
$$dP = \underbrace{dA}_{(1)} \cos\theta \underbrace{d\omega}_{(2)} \underbrace{L}_{(3)}$$



$$dw: d\theta d\phi \text{ (horiz)}$$

$$dw: d\theta d\phi \sin\theta$$

$$dP = dA \cos\theta dw L$$



$$dP = dA \cos\theta d\theta d\phi \sin\theta L$$

want power per unit Area
($dA=1$)

$$dE = L \cos\theta \sin\theta d\phi d\theta$$

for total E, integrate over ϕ, θ

$$E = \int_{\theta=0}^{\pi/2} \int_{\phi=0}^{2\pi} L \cos\theta \sin\theta d\phi d\theta$$

not dependent on direction (isotropic)

$$E = L \int \dots$$

$$E = L \int_{\theta=0}^{\pi/2} \left[\cos\theta \sin\theta \phi \right]_0^{2\pi} d\theta$$

$$\cos\theta \sin\theta \cdot 2\pi$$

$$E = 2\pi L \int_{\theta=0}^{\pi/2} \underbrace{\cos\theta \sin\theta}_{d\theta} d\theta$$

$$\sin \theta \cos \theta = \frac{1}{2} \sin 2\theta$$

$$E = 2\pi L \int_{\theta=0}^{\pi/2} \frac{1}{2} \sin 2\theta \, d\theta$$

$$2\pi L \cdot \left[-\frac{1}{4} \cos 2\theta \right]_0^{\pi/2}$$

$\frac{1}{4} - -\frac{1}{4}$

E
irradiance
M
exitance
emittance

$$2\pi L \cdot \frac{1}{2}$$

$$\boxed{E = \pi L}$$

↓↓↓↓

$$\boxed{M = \pi L}$$

↑↑↑

IRIS : Infra Red Imaging Subsystem
AUSTRALIS-1

700-1100 nm

energy output of sun

consider sun: black body radiator

$$T = 5800^{\circ} \text{K}$$

ϵ : 0.99 emissivity

R_s : radius $6.96 \times 10^8 \text{ m}$

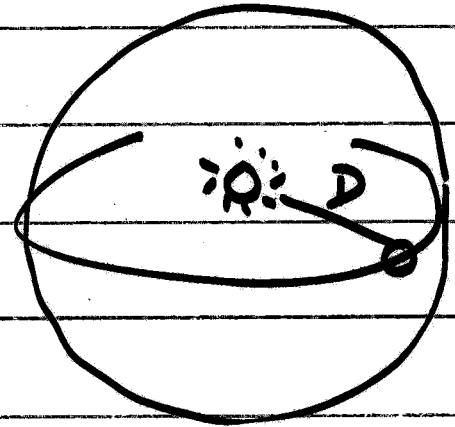
D : earth $1.496 \times 10^{11} \text{ m}$

σ : stefan boltzmann
constant

$$5.67 \times 10^{-8} \text{ W m}^{-2} \text{ K}^{-4}$$

$$\begin{aligned}\Phi_{\text{sun}} &= 4\pi R_s^2 T^4 \sigma \epsilon \\ &= 3.87 \times 10^{26} \text{ W (J/s)}\end{aligned}$$

all wavelengths
all directions



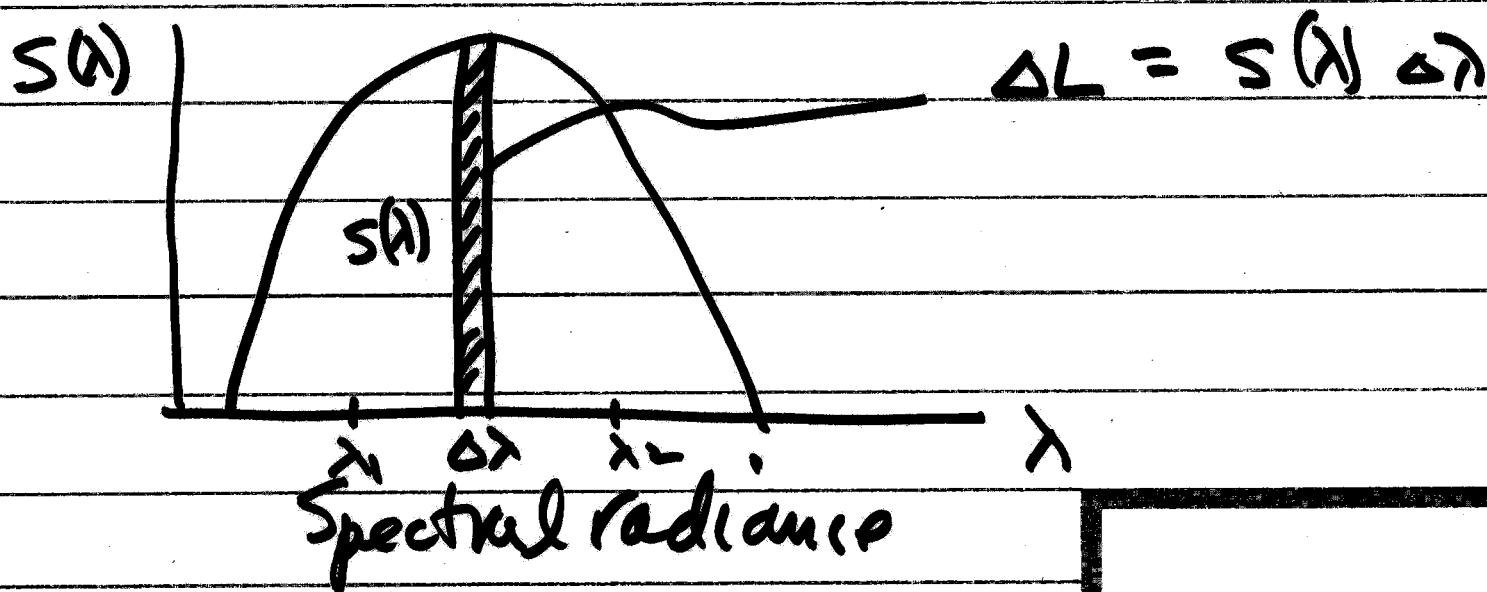
$$E_e = \frac{\Phi_{\text{sun}}}{A} = \frac{\Phi_{\text{sun}}}{4\pi D^2}$$

$$E_e = 1.376 \times 10^3 \text{ W/m}^2$$

irradiance @ top of
atmosphere

$1.376 \times 10^3 \text{ W/m}^2$ all wavelengths

IRIS : 700 - 1100 nm



$$L_{\lambda_1, \lambda_2} = \sum_{\lambda_i = \lambda_1}^{\lambda_2} S(\lambda_i) \cdot \Delta \lambda$$

$$L_{\lambda_1, \lambda_2} = \int_{\lambda_1}^{\lambda_2} S(\lambda) d\lambda$$

$$S(\lambda) = \frac{2hc^2}{\lambda^5 (e^{hc/\lambda kT} - 1)}$$

planck's formula

h : ~~6.26~~ 6.626×10^{-34} Js planck

k : 1.3807×10^{-23} JK⁻¹ boltzmann

c : 3×10^8 m/s

T : temp °K

λ : wavelength (m)

I provide Matlab code to numerically integrate

$$M = \pi L \cdot \epsilon$$

↑
integration

$$\phi = 4\pi R_s^2 M$$

$$E_e = \frac{\phi}{4\pi D^2}$$

now be $< 1376 \text{ w/m}^2$

$$\underline{\underline{E_e = 377.4 \text{ w/m}^2}}$$

reflectivity or albedo (visible)

forest ~ 10%

crops ~ 5-15%

grass ~ 5-30%

soil ~ 5-30%

sand ~ 20-40%

ice ~ 30-40%

old snow ~ 50-70%

new snow ~ 75-90%

design maximum

$$\underline{\underline{0.7 = \gamma_d}}$$

$$E_e = 377.4 \text{ W/m}^2 \quad \text{restricted to NIR}$$

0.7 albedo

$$M_e = E_e \cdot \alpha_d = 264.2 \text{ W/m}^2$$