

# I. Derivation of the Equations of Motion.

## A. Review of Cartesian Tensor Notation

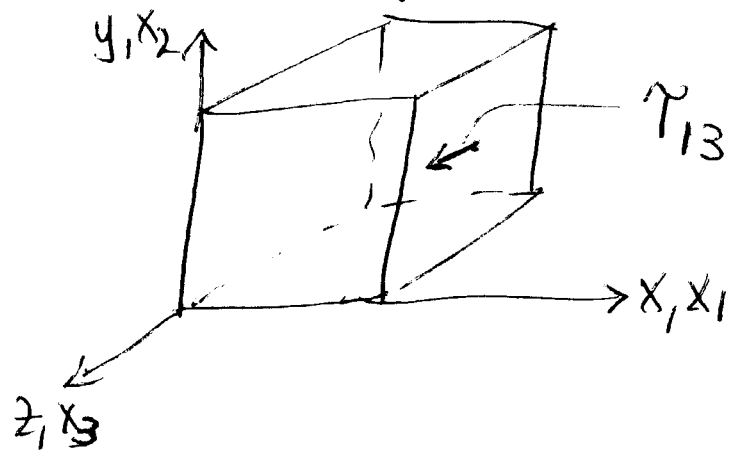
References for Cartesian Tensors:

Botkov, Chap 16; Jeffreys and Jeffreys, Chapter 3<sup>2nd</sup>

Fung, Foundations of Solid Mechanics, Chapter 2; etc, etc.

### 1. Introduction: Shear Stress: (Tensor)

Use the symbol  $\tau_{ij}$  for the shear stress on the  $i$  face of a box, in the  $j$ 'th direction:



Sign convention

$\tau_{ij} > 0$  if pointing in  $+j$  direction on  $+i$  face, or in  $-j$  direction on  $-i$  face.

Note: for an infinitesimal cube (shear stress at a point),

$$\tau_{ij} = \tau_{ji}$$

$$\underline{\tau} = \tau_{ij} = \begin{bmatrix} \tau_{11} & \tau_{12} & \tau_{13} \\ \tau_{21} & \tau_{22} & \tau_{23} \\ \tau_{31} & \tau_{32} & \tau_{33} \end{bmatrix}$$

-(check of transposed)

Shear is a tensor, not just a collection of 9 numbers; it must be invariant under changes in the coordinate system (rotations, translations, etc).

## 2. The Summation Convention.

The divergence of the velocity  $y$  can be written as

$$\nabla \cdot y = \text{div } y = \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} + \frac{\partial w}{\partial z}$$

$$= \frac{\partial u_1}{\partial x_1} + \frac{\partial u_2}{\partial x_2} + \frac{\partial u_3}{\partial x_3}$$

$$= \sum_j \frac{\partial u_j}{\partial x_j} = \frac{\partial u_j}{\partial x_j}$$

with the  $\sum$  understood, whenever an index is repeated (Einstein summ. conv.)

Note also that  $\underline{a} \cdot \underline{b} = a_1 b_1 + a_2 b_2 + a_3 b_3 = a_j b_j$   
 where  $\underline{a} = a_1 \hat{e}_x + a_2 \hat{e}_y + a_3 \hat{e}_z = (a_1, a_2, a_3)$

## 3. Some special tensors:

- $\delta_{ij}$  - the Kronecker delta function.  $= 1$  if  $i=j$   
 (or identity matrix)  $= 0$  if  $i \neq j$
- $\epsilon_{ijk}$  - the permutation tensor  $= 1$  if  $ijk$  follow in cyclical order (e.g. 12312...)  
 $= -1$  if out of order by 1 permutation (e.g. 213210...)  
 $= 0$  otherwise (any repeated #)

See J&J p. 69ff

Note:  $\delta_{ij} \delta_{jk} = \delta_{ik}$ , since  $\delta_{ij} a_j = a_k$

why? ① if  $i \neq j$  or  $j \neq k$ ,  $= 0$ ; if  $i = j$  and  $j = k$ ,  $= 1$  but  $i = k$ .

②  $\delta_{i1} \delta_{1k} + \delta_{i2} \delta_{2k} + \delta_{i3} \delta_{3k} = \delta_{ij} \delta_{jk}$ ; if  $i \neq k$  all  $= 0$   
if  $i = k$ , are  $= 1$

Note:  $\epsilon_{ijk} \epsilon_{ilm} = \delta_{jl} \delta_{km} - \delta_{jm} \delta_{kl}$  very useful

why? terms with  $i = j$  or  $i = k = 0$ ,  $i = l$  or  $i = m = 0$

① if  $j = k$  or  $l = m$ ,  $= 0$ .

if  $j$  is one above  $i$ , then  $l$  must be above or below, otherwise term is zero. If  $j$  is above  $i$ , then  $k$  is above  $j$  (only possibility). If  $l$  is below  $i$ , then  $m$  must be below  $l$  (only possibility)

Thus, if in same order, get  $+1 \cdot +1$  or  $-1 \cdot -1 = +1$ .

if in different order, get  $-1 \cdot +1$  or  $+1 \cdot -1 = -1$ .

same order means when  $j = l$  and  $k = m$ ; different order

means when  $j = m$  and  $k = l$   
explains the sign argument.  
use "same order" or "different order" argument.

(2) other way: if  $i=1, j=2$  and  $k=3$  or  $j=3$  and  $k=2$ , otherwise  $\epsilon_{ijk}=0$ . Similar with 2d term.

#### 4. Vector Calculus in Index Notation.

Note  $\underline{a} \cdot \underline{b} = a_j b_j$  as before.

$$\begin{aligned} \text{Note } \underline{a} \times \underline{b} &= \epsilon_{ijk} a_j b_k = (a_2 b_3 - a_3 b_2) \hat{e}_1 \\ &\quad + (a_3 b_1 - a_1 b_3) \hat{e}_2 + (a_1 b_2 - a_2 b_1) \hat{e}_3 \\ &= \begin{vmatrix} \hat{e}_1 & \hat{e}_2 & \hat{e}_3 \\ a_1 & a_2 & a_3 \\ b_1 & b_2 & b_3 \end{vmatrix} \end{aligned}$$

$$\text{Thus, } \nabla \times \underline{y} = \text{curl } \underline{y} = \epsilon_{ijk} \frac{\partial}{\partial x_j} u_k = \begin{vmatrix} \hat{e}_1 & \hat{e}_2 & \hat{e}_3 \\ \frac{\partial}{\partial x_1} & \frac{\partial}{\partial x_2} & \frac{\partial}{\partial x_3} \\ u_1 & u_2 & u_3 \end{vmatrix}$$

Note: Compute  $\nabla \times (\nabla \times \underline{y}) = \nabla \times \underline{w}$ :

$$\begin{aligned} \nabla \times \nabla \times \underline{y} &= \epsilon_{ijk} \frac{\partial}{\partial x_j} \left[ \epsilon_{k\ell m} \frac{\partial}{\partial x_\ell} u_m \right] = \epsilon_{ijk} \epsilon_{k\ell m} \frac{\partial}{\partial x_j} \frac{\partial}{\partial x_\ell} u_m \\ &= \epsilon_{kij} \epsilon_{k\ell m} \frac{\partial}{\partial x_j} \left( \frac{\partial}{\partial x_\ell} u_m \right) = (\delta_{j\ell} \delta_{im} - \delta_{jm} \delta_{i\ell}) \frac{\partial^2 u_m}{\partial x_j \partial x_\ell} \\ &= \frac{\partial}{\partial x_i} \left( \frac{\partial u_j}{\partial x_j} \right) - \frac{\partial^2}{\partial x_j \partial x_j} (u_i) = \nabla (\nabla \cdot \underline{y}) - \nabla^2 (\underline{y}) \\ &= \nabla (\nabla \cdot \underline{y}) - \nabla \cdot (\nabla \underline{y}) \end{aligned}$$

to here 10-10

(5)

## B. Some Theorems from Vector Calculus

(1) Stokes Theorem:

$$\Gamma = \int_C \underline{y} \cdot d\underline{x} = \iint_S (\text{curl } \underline{y}) \cdot \hat{n} \, dS$$

(curl is local 'circulation')

The line integral around a curve is equal to the surface integral of the curl. (remember right-hand rule for calculating line integral, and for curl)

(2) Gauss' Theorem:

$$\iint_S \underline{y} \cdot \hat{n} \, dS = \text{flux out} = \iiint_V \text{div } \underline{y} \, dV$$

(divergence is the local 'flux out')

Note: There is a form of Gauss' theorem which is sometimes useful when working with the curl.

$$\iiint_V (\nabla \times \underline{f}) \, dV = \iint_S (\hat{n} \times \underline{f}) \, dS$$

This derivation is left for a homework problem.  
(a lemma to Gauss' theorem)

Note also: These theorems are also valid for tensors, not just for vectors:

$$\iint_S \underline{\underline{T}} \cdot \hat{n} \, dS = \iiint_V \text{div} \underline{\underline{T}} \, dV$$

$$\underline{\underline{T}} \cdot \hat{n} = \tau_{ij} n_j, \quad \text{div} \underline{\underline{T}} = \frac{\partial}{\partial x_j} (\tau_{ij})$$

## C. Review of Thermodynamics

① First Law of Thermodynamics - Conservation of Energy.

If  $\Delta Q$  is the heat added to a system, and  $\Delta W$  is the work done on a system, and if  $E$  is the internal energy of a system,

$$\Delta W + \Delta Q = \Delta E, \text{ or}$$

$$\boxed{dW + dQ = dE}$$

Cannot be integrated without knowing path

( $d$  - not a perfect differential)

② Second Law of Thermodynamics - Entropy

Entropy is a variable of state, and for a closed system,  $S$  increases in any spontaneous process. Also, for a reversible process,

$$S_B - S_A = \int_A^B \frac{dQ}{T}, \text{ where } A \rightarrow B \text{ is a reversible process.}$$

$$\text{then } ds|_{\text{rev.}} = \frac{dQ}{T}|_{\text{rev.}}$$

or  $Tds = dE + PdV$ , where we have set  $dW = PdV$ , assuming all work done through pressure.  
Thermodynamic Identity

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\* ~~Pressure forces are not true~~ <sup>Not ~~still~~ true unless</sup> all of the work is done through pressure forces; but shear usually produces heat, not work - see Schilling - Eegderiv.

This equation involves only variables of state, so it can be integrated, and the result is path-independent

$$\int_1^2 ds = S_2 - S_1 = \int_1^2 \frac{dE}{T} + \int_1^2 \frac{P}{T} dV$$

Since the integral of S is path independent, ds is a perfect differential, and

$$ds = \frac{1}{T} dE + \frac{P}{T} dV = \frac{\partial S}{\partial E}|_V dE + \frac{\partial S}{\partial V}|_E dV$$

$$\text{thus, } \frac{1}{T} = \frac{\partial S}{\partial E}|_V, \frac{P}{T} = \frac{\partial S}{\partial V}|_E, \text{ etc.}$$

Note: Can get various relations by equating the mixed partials, e.g.

$$\frac{\partial^2 S}{\partial E \partial V} = \frac{\partial^2 S}{\partial V \partial E} \Rightarrow \frac{\partial}{\partial V} \left( \frac{1}{T} \right) = \frac{\partial}{\partial E} \left( \frac{P}{T} \right)$$

$$-\frac{1}{T^2} \frac{\partial T}{\partial V} \Big|_E = \frac{T \frac{\partial P}{\partial E} - P \frac{\partial T}{\partial E}}{T^2} \quad (E, V \text{ space})$$

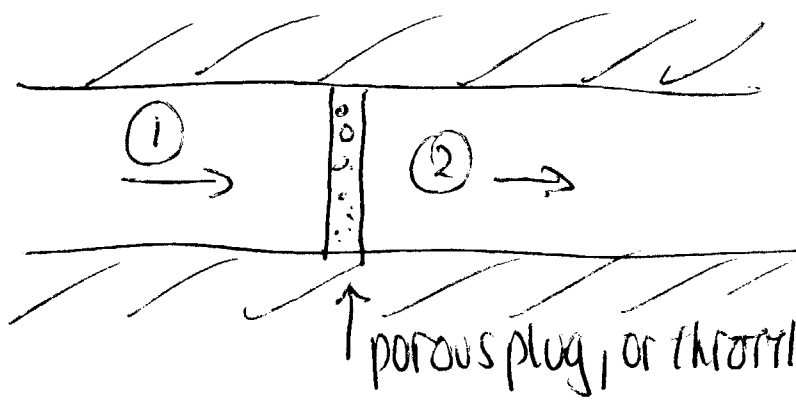
$$\frac{\partial T}{\partial V} \Big|_E = P \frac{\partial T}{\partial E} \Big|_{\cancel{V}} - T \frac{\partial P}{\partial E} \Big|_{\cancel{V}} \quad (\text{get Maxwell relations this way})$$

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### ③ Example - Joule-Thomson Throttling Process

See  
LAR.



Steady flow through a heat-insulating pipe with a porous plug.

Assume fluid flows slowly, so that its kinetic energy is small compared to its enthalpy ( $\frac{1}{2} \rho u^2 \ll h$ )

Consider passage of a unit mass through the plug:

825



9

Change in internal energy is

$$\Delta E = e_2 - e_1$$

work done (as fluid leaves left and pushes out on right) is

$$W = \int_1 p dv \approx \int_2 p dv$$

$P_1$  and  $P_2$  are constant in this steady process, so

$$\text{work done on gas} = W = P_1 v_1 - P_2 v_2 \quad (\text{goes out left, comes in on right})$$

where  $v_1$  is volume of gas when all on l.h.s.

$v_2$  is volume of gas when all on r.h.s.

By the first law, since the heat flux out is zero, (assume no conduction in fluid or to walls)

$$\Delta E = W$$

$$e_2 - e_1 = P_1 v_1 - P_2 v_2$$

$$h_1 = e_1 + P_1 v_1 = e_2 + P_2 v_2 = h_2 \quad (h = e + pv)$$

In adiabatic flow through a resistance, the enthalpy per unit mass is the same upstream and downstream.

"adiabatic" - a system without heat conduction to the outside.

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Gottlieb  
Fri 11-2  
Dontskip  
EATS.

ep blunt nose in S.S. flow →

(4) Perfect gas:  $pV = RT$   $R$  is the specific gas constant

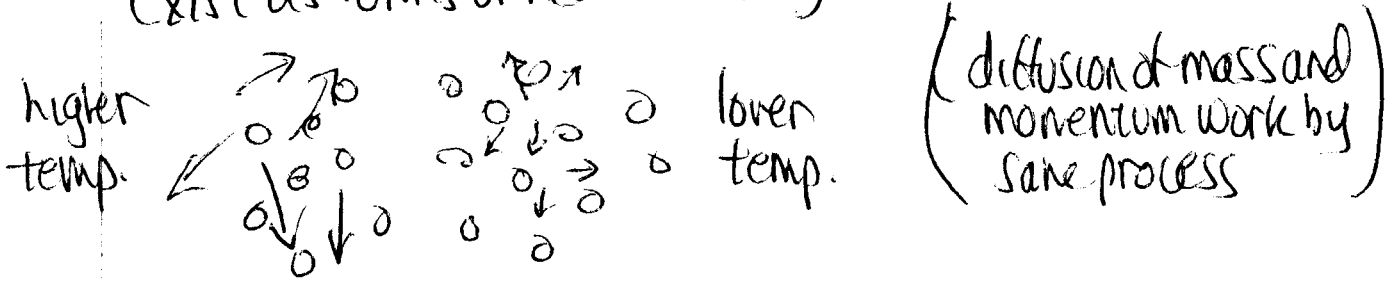
calorically perfect:  $h = c_p T$ ,  $e = c_v T$ ,  $c_p$  &  $c_v$  are constant,  $T$  absolute.

will use to save on laborious calculations.

### D. Heat Transfer (Conduction) (Diffusion)

Diffusion is transport by molecular motion.

The simplest example is conductive heat transfer (on the molecular scale, only conduction and radiation exist as forms of heat transfer).



more molecules from higher temp randomly move to lower temp than vice versa. Thus, high-speed molecular motion region diffuses into low-speed region. Heat diffuses.

Diffusion is approximated as a linear process - heat diffuses down the temperature gradient.

thus, say

$$\vec{q} = \vec{q}(\nabla T, \text{second. deriv. of } T) \quad \begin{matrix} \nearrow \text{ignore} \\ \text{assume } \nabla T \text{ small} \\ \text{(compared to } \rho \text{ etc?)} \end{matrix}$$

assume that  $\vec{q}$  is a linear vector function of  $\nabla T$ .

$$q_i = a_{ij} \frac{\partial T}{\partial x_j} = \vec{a} \cdot \nabla T$$

$\vec{a}$  is the heat conductivity tensor

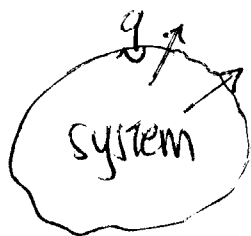
for an isotropic substance, all directions are the same,

$$a_{ij} = -\lambda \delta_{ij}$$

thus,  $\vec{q} = -\lambda \nabla T$  where  $\lambda$  is some constant.

usually use  $\vec{q} = -k \nabla T$   $k$  is the thermal conductivity. (see white p 30)

Now consider the flow of heat in some substance.



$$\Delta(\text{Total heat enclosed}) = \int_S \vec{q} \cdot \hat{n} \Delta t$$

$$\frac{\partial}{\partial t} \left( \int_V \rho e dV \right) = - \int_S \vec{q} \cdot \hat{n} dS$$

For a discussion of the limitations on Fourier's Law for heat conduction, see "Heat Under the Microscope", by Maasilta and Minnich, Physics Today, August 2014, pp. 27-32. When the gradients in temperature become significant on the scale of the mean free path of the phonons, Fourier's law breaks down.

for perfect gas,  $e = c_v T$

$$\iiint_V \frac{\partial}{\partial t} (e c_v T) dV + \iiint_V \text{div } \underline{q} dV = 0$$

using Gauss' law  
and interchange  
 $\int \frac{\partial \phi}{\partial t}$

true for any  $V$ , so

$$\frac{\partial}{\partial t} (e c_v T) + \text{div } \underline{q} = 0 \text{ everywhere}$$

assume  $e$  constant, then, using  $\underline{q} = -k \nabla T$ ,  $k$  constant,

$$e c_v \frac{\partial T}{\partial t} - \nabla \cdot k \nabla T = 0$$

$$\frac{\partial T}{\partial t} = \frac{k}{e c_v} \nabla^2 T \quad (\nabla^2 T \equiv \nabla \cdot \nabla T)$$

$k/e c_v = \alpha = \text{thermal diffusivity}$  - White p. 83 -  
(but see p. 34 ~~different~~ <sup>typo</sup>).  
See Thompson problem 6.13

thus,  $\boxed{\frac{\partial T}{\partial t} = \alpha \nabla^2 T - \text{the heat equation.}}$

for the conduction of heat  
in a stationary fluid.

• mom. diffusion more complex  
eqn, since mom. a vector;  
but same process.

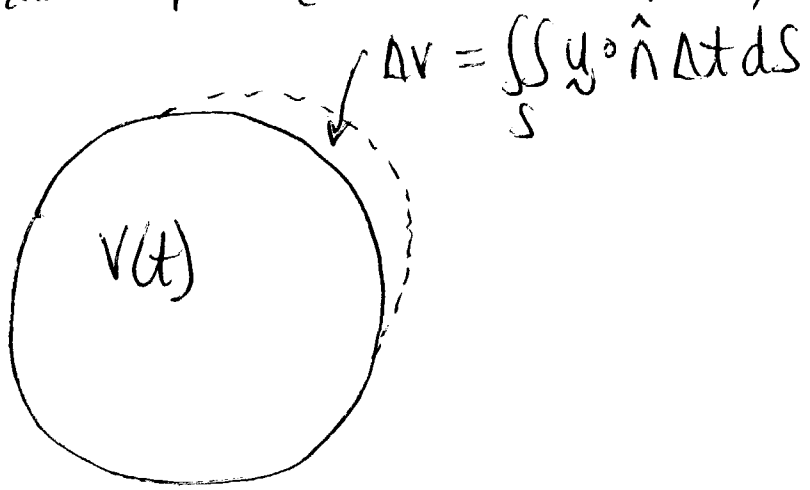
• Same equation for  
mass diffusion

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### E. The Reynolds Transport Theorem

Want to write laws for conservation of mass, momentum, and energy. But these laws are usually written for a fixed mass; need to rephrase them for volumes fixed in space (Eulerian description)

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Consider any quantity  $\phi$  - could be scalar, vector, tensor (2nd order tensor). Let  $\Phi$  be the total amount of  $\phi$  in some volume,

$$\Phi = \iiint_{V(t)} \phi dV \quad \text{where } V \text{ moves with the fluid}$$

What is the rate of change of  $\Phi$ ?

$$\frac{d\Phi}{dt} = \lim_{\Delta t \rightarrow 0} \frac{\Phi(t + \Delta t) - \Phi(t)}{\Delta t}$$

$\Phi$  changes for 2 reasons: ① volume enclosed by moving fluid changes w/ time ② fluid inside the volume changes even if remains inside

$$\Phi(t+\Delta t) = \iiint_{V(t+\Delta t)} \phi(t+\Delta t) dV$$

$\swarrow$   
 $\phi(x, y, z, t+\Delta t)$

$$\Phi(t) = \iiint_{V(t)} \phi(t) dV$$

$$\Phi(t+\Delta t) - \Phi(t) = \iiint_{V(t+\Delta t)} \phi(t+\Delta t) dV - \iiint_{V(t)} \phi(t) dV$$

Now, add & subtract  $\iiint_V \phi(t+\Delta t) dV$

$$\Phi(t+\Delta t) - \Phi(t) = \underbrace{\iiint_{V(t+\Delta t)} \phi(t+\Delta t) dV - \iiint_{V(t)} \phi(t+\Delta t) dV}_{\text{I}}$$

$$+ \underbrace{\iiint_{V(t)} [\phi(t+\Delta t) - \phi(t)] dV}_{\text{II}}$$

What is  $\mathcal{I}$ ?  $V(t+\Delta t) = V(t) + \Delta V \Rightarrow$

$$\mathcal{I} = \iiint_{V(t+\Delta t) - V(t)} \phi(t+\Delta t) dV = \iint_S \phi y \cdot \hat{n} \Delta t dS$$

$\uparrow$  distance  $L$  to surface  
 $\uparrow$  swept carried w/ fluid which moved out.

dividing both sides of equation by  $\Delta t$ , we have that

$$\frac{\Phi(t+\Delta t) - \Phi(t)}{\Delta t} = \iint_S \phi y \cdot \hat{n} dS + \iiint_V \frac{\phi(t+\Delta t) - \phi(t)}{\Delta t} dV$$

as  $\Delta t \rightarrow 0$ , we get, in the limit,

$\frac{d}{dt} \int \phi dV$   
following fluid

$$\left. \frac{d\Phi}{dt} \right|_{\text{following the fluid}} = \iint_S \phi y \cdot \hat{n} dS + \iiint_V \frac{\partial \phi}{\partial t} dV$$

$\nwarrow$  fixed in space  $\rightarrow V$

First Form of Reynolds Transport Theorem.

Use Gauss' Theorem to change the surface integral to a volume  $\int$

Second Form.

$$\left. \frac{d\Phi}{dt} \right|_{\text{following the fluid}} = \iiint_V (\text{div}(\phi y) + \frac{\partial \phi}{\partial t}) dV = \frac{d}{dt} \left| \iiint_V \phi dV \right|_{\text{following fluid}}$$

$\nwarrow$  fixed in space

Remember,  $\phi$  can be anything which is carried with the fluid.

### F<sub>0</sub> The Continuity Equation

The simplest equation of motion expresses mass conservation. Since mass is neither created nor destroyed (if we don't consider nuclear reactions) then the amount of mass within a volume of fluid that stays with the fluid particles must remain the same. That is,

$$\frac{d}{dt} \int_{\text{following fluid}} \rho \, dV = 0 = \int_{\text{fixed in space}} \left[ \frac{\partial \rho}{\partial t} + \text{div}(\rho \mathbf{u}) \right] dV$$

where I have used Reynolds transport theorem.

Since this ~~must be~~<sup>is</sup> true for an arbitrary volume  $V$ , it must be true at every point (from 'weak' to 'strong' forms - actually 'weak', averaging form, is more physically true - important for math. proofs). Thus

$$\frac{\partial \rho}{\partial t} + \text{div}(\rho \mathbf{u}) = 0$$

Mass conservation or Continuity Equation



Of course, we can express this same equation in various 'control volume' or integral forms, using Gauss' theorem;

e.g. 
$$\iiint_V \frac{\partial \rho}{\partial t} dv + \iint_S \rho u \cdot \hat{n} ds = 0$$

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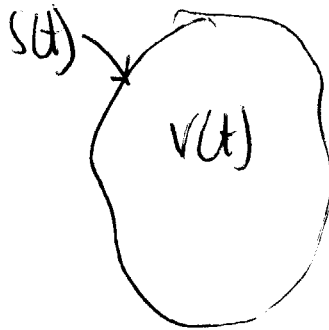
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Go Momentum Equation

We know that

$$\underline{F} = \frac{d}{dt}(\text{momentum})$$

Consider the fluid within some volume  $V$ , which moves with the fluid.



We must have that

$$\underline{\Sigma F} = \frac{d}{dt} \iiint_{V(t)} \rho \underline{y} dv$$

↑ moving with the fluid.

Or, using Reynolds transport theorem,

$$\Sigma F = \iiint_V \frac{\partial}{\partial t}(\rho y) dv + \iint_S \rho y (y \cdot \hat{n}) ds$$

with the volume and  $S$  now fixed in space.

Use the extended form of Gauss Law,

$$\Sigma F = \iiint_V \frac{\partial}{\partial t}(\rho y) dv + \iint_V \operatorname{div}(\rho y y) d$$

$$(\operatorname{div}(\rho y y) = \frac{\partial}{\partial x_j}(\rho u_i u_j))$$

$$\text{or } \Sigma F = \iiint_V \left[ \frac{\partial}{\partial t}(\rho y) + \operatorname{div}(\rho y y) \right] dv$$

What are the forces  $\Sigma F$ ? Two types:

(1) Surface forces: act on contact, on the surface  
(pressure, shear)

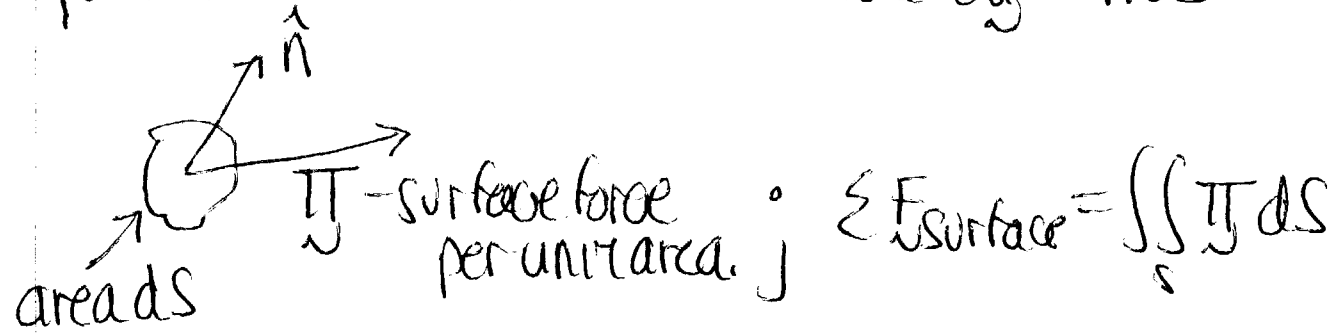
(2) Body forces: Act at a distance, on fluid in whole volume.  
(gravity, electromagnetic forces, etc)

If we assume that the body forces are independent of the velocity, and so on (like gravity), then

$$\sum \underline{F}_{\text{body}} = \iiint_V \underline{f} \, dV \quad \text{where } \underline{f} \text{ is the } \underline{\text{body force}} \underline{\text{per unit volume.}}$$

I won't do E & M forces here. (does anybody want them?)

The surface forces usually considered are the pressure and the shear (are there others? - surface tension?)  
 Combine these into one symbol, let  $\underline{\pi}$  be the force per unit area on an element of surface  $d\mathbf{a} = \hat{n} \, dS$



$\underline{\pi}$  depends on the orientation of the surface  $dS$  (depends on  $\hat{n}$ ) as well as on its position; it is thus more complex than  $\underline{f}$ .

We will return to this. For now, add these

$$\sum \underline{F} = \sum \underline{F}_{\text{body}} + \sum \underline{F}_{\text{surface}} = \iiint_V \underline{f} \, dV + \iint_S \underline{\pi} \, dS$$

So that the momentum conservation equation is

$$\int_V \left[ \frac{\partial}{\partial t} (\rho y) + \text{div} (\rho y u) \right] dV = \int_V f dV + \int_S \Pi dS$$

where all integrals are now fixed in space.

Need to simplify this equation to use it. What is  $\Pi$ ?

Clearly,  $\Pi$  can depend on the surface orientation (think of how pressure force does).  $\Pi$  is thus a function of  $\hat{n}$ . What kind of function?

To avoid infinite forces on a closed volume as the volume shrinks to zero (no  $\infty$  volume forces at a point), must have  $\Pi$  a linear function of  $\hat{n}$ .

(see Batchelor p. 9 & 10, etc) Proof left as exercise. -hand out-

8-29-94

If  $\Pi_i$  is to be a linear function of  $n_j$ , the most complex possible is

$$\Pi_i = \sigma_{ij} n_j$$

$$\underline{\Pi} = \underline{\sigma} \cdot \hat{n}$$

$\sigma_{ij}$  the stress tensor (because has to be invariant under rotations of coords, etc)

Now this  $\tau_{ij}$  includes pressure forces and shear forces, for the moment — it represents all the surface forces.

Then we can rewrite momentum equation as:

$$\iiint_V \left[ \frac{\partial}{\partial t}(\rho \mathbf{y}) + \operatorname{div}(\rho \mathbf{y} \mathbf{y}) \right] dV = \iiint_V \underline{f} dV + \iint_S \underline{\sigma} \cdot \hat{\mathbf{n}} dS$$

use Gauss's law on last term:

$$\iiint_V \left[ \frac{\partial}{\partial t}(\rho \mathbf{y}) + \operatorname{div}(\rho \mathbf{y} \mathbf{y}) \right] dV = \iiint_V \underline{f} dV + \iiint_V \operatorname{div} \underline{\sigma} dV$$

or, combining all volume  $V$ ,

$$\iiint_V \left[ \frac{\partial}{\partial t}(\rho \mathbf{y}) + \operatorname{div}(\rho \mathbf{y} \mathbf{y}) - \underline{f} - \operatorname{div} \underline{\sigma} \right] dV = 0$$

Since this is true for an arbitrary volume  $V$ , it must be true everywhere (weak soln  $\rightarrow$  strong soln)

$$\frac{\partial}{\partial t}(\rho \mathbf{y}) + \operatorname{div}(\rho \mathbf{y} \mathbf{y}) - \underline{f} - \operatorname{div} \underline{\sigma} = 0$$

or

$$\frac{\partial}{\partial t}(\rho \mathbf{y}) + \operatorname{div}(\rho \mathbf{y} \mathbf{y} - \underline{\sigma}) = \underline{f}$$

Now, separate out the thermodynamic pressure from the rest of the surface stress

$$\underline{\underline{\sigma}} = \underline{\underline{\tau}}_{ij} = -\delta_{ij} p + \underline{\underline{\gamma}}_{ij} = -\underline{\underline{I}} p + \underline{\underline{\gamma}}$$

where  $\underline{\underline{\gamma}}$  is the viscous shear part (note  $\tau_{ii} \neq 0$  in general)  
( $p$  is pressure inside) (was force on element)

Then we have

$$\frac{\partial}{\partial t}(\rho \underline{y}) + \text{div}(\rho \underline{y} \underline{y} + p \underline{\underline{I}} - \underline{\underline{\gamma}}) = \underline{f}$$

here 8/26/16.

hard term to deal with.

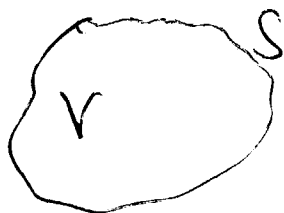
$\underline{\underline{\gamma}}$  can depend only on the derivatives of  $\underline{y}$ , and not on  $\underline{y}$  (must be Galilean invariant) We will approximate that

$\underline{\underline{\gamma}}$  is linear in the first derivatives of  $\underline{y}$  (Newtonian fluid).

Just like approximating  $q$  linear in <sup>1st</sup> derivatives of  $T$  (Fourier heat conduction).

here  $\tau_{ii}$  H9

### H0 Energy Equation



Look at energy content of fluid inside a volume which moves with the fluid.

actually first term in expansion of  $\frac{E_0}{(1-v^2/c^2)^{1/2}}$  French p. 22

non-relativistic

$$\frac{d}{dt} \int_V (\rho e + \frac{1}{2} \rho u^2) dV = \text{Rate of (work done + heat added)}$$

moving w/ fluid  $\uparrow$   $V(t)$   
 $\uparrow$  internal energy  $\uparrow$  KE

( $e =$  internal energy per unit mass)  
( $u^2 \equiv u \cdot u$ )

use Reynolds transport theorem:

$$\frac{d}{dt} \int_{V(t)} \left[ \frac{\partial}{\partial t} (\rho e + \frac{1}{2} \rho u^2) + \text{div} [\underline{y} (\rho e + \frac{1}{2} \rho u^2)] \right] dV = \text{Rate of (work + heat)}$$

Forces:  $\underline{\pi}$  on  $S$ ,  $\underline{f}$  on  $V$  (for work done)

heat added:  $-\underline{q} \cdot \underline{\hat{n}} dA$  ( $\underline{y} = \lim \frac{\Delta \underline{x}}{\Delta t}, \Delta t \rightarrow 0$ )

$$\text{rate of work done by } \underline{\pi} = \iint_S \underline{y} \cdot \underline{\pi} dS = \iint_S \underline{y} \cdot \underline{q} \cdot \underline{\hat{n}} dS = \iiint_V \text{div} (\underline{y} \cdot \underline{q}) dV$$

$$\text{rate of work done by } \underline{f} = \iiint_V \underline{y} \cdot \underline{f} dV$$

$$\text{rate of heat added} = -\iint_S \underline{q} \cdot \underline{\hat{n}} dS = -\iiint_V \text{div } \underline{q} dV$$

so whole equation becomes:

$$\iiint_{V \text{ fixed}} \left[ \frac{\partial}{\partial t} (\rho e + \frac{1}{2} \rho u^2) + \text{div} (y (\rho e + \frac{1}{2} \rho u^2)) \right] dV$$

$$= \iiint_V (\text{div} (y \cdot \underline{\underline{\tau}} - q) + y \cdot \underline{\underline{f}}) dV$$

need to  
 (29) (30) (31) (32)  
 for this

Since this is true for arbitrary volume, true everywhere:

$$\frac{\partial}{\partial t} [\rho (e + \frac{1}{2} u^2)] + \text{div} (y (\rho e + \frac{1}{2} \rho u^2) + q - \underline{\underline{\tau}} \cdot y) - \underline{\underline{f}} \cdot y = 0$$

$$\text{but } \underline{\underline{\tau}} = -p \underline{\underline{I}} + \underline{\underline{\tau}} \cdot y$$

$$\underline{\underline{\tau}} \cdot y = -p \underline{\underline{I}} \cdot y + \underline{\underline{\tau}} \cdot y$$

$$= -p y + \underline{\underline{\tau}} \cdot y$$

so

$$\frac{\partial}{\partial t} [\rho (e + \frac{1}{2} u^2)] + \text{div} [\rho y (e + \frac{1}{2} u^2) + q + p y - \underline{\underline{\tau}} \cdot y] - \underline{\underline{f}} \cdot y = 0$$

$$\frac{\partial}{\partial t} [\rho (e + \frac{1}{2} u^2)] + \text{div} [\rho y (e + \frac{1}{2} u^2 + \frac{p}{\rho}) + q - \underline{\underline{\tau}} \cdot y] - \underline{\underline{f}} \cdot y = 0$$

$$\text{but } e + p/\rho = e + p v = h$$

$$\frac{\partial}{\partial t} [\rho (e + \frac{1}{2} u^2)] + \text{div} [\rho y (h + \frac{1}{2} u^2) + q - \underline{\underline{\tau}} \cdot y] - \underline{\underline{f}} \cdot y = 0$$

$$\frac{\partial}{\partial t} [\rho (e + \frac{1}{2} u^2)] + \text{div} [\rho y (h + \frac{1}{2} u^2) + q - \underline{\underline{\tau}} \cdot y] = \underline{\underline{f}} \cdot y$$

Energy equation (scalar)

— can be simplified to heat conduction eqn. —



8-31-94

8-27-

Eqs so  
for general,  
no E&M,  
relativity...

(25)

# I. Fluid Characteristics

Equations of Motion are:

$$\frac{\partial \rho}{\partial t} + \text{div}(\rho \mathbf{u}) = 0 \quad (\text{continuity})$$

$$\frac{\partial(\rho \mathbf{u})}{\partial t} + \text{div}(\rho \mathbf{u} \mathbf{u} + \mathbf{P} \mathbf{I} - \mathbf{T}) = \mathbf{f} \quad (\text{momentum})$$

$$\frac{\partial(\rho e + \frac{1}{2} \rho u^2)}{\partial t} + \text{div}[\rho \mathbf{u} (h + \frac{1}{2} u^2) - \mathbf{T} \cdot \mathbf{u} + \mathbf{q}] = \mathbf{f} \cdot \mathbf{u} \quad (\text{energy})$$

Too general - Need equations for fluid characteristics (eqns of state) for  $\mathbf{T}$ ,  $\mathbf{q}$ , etc.

For gases, some of this can be derived from kinetic theory. For liquids, this has not been successful (long-short range coupling, multi-body interactions, same problem as in turbulence). Use phenomenological models for liquids; might as well use for gases here too.

①  $\mathbf{q}$ , the heat flux vector, depends on spatial derivatives of  $T$ . Assume depends on 1st derivatives only -  
Fourier's Law:  $\mathbf{q}$  is a linear function of  $\mathbf{1}^{\text{st}}$  deriv. of  $T$ .

Thus,  $\underline{q} = \underline{\kappa} \cdot \nabla T$  or  $q_i = \kappa_{ij} \frac{\partial T}{\partial x_j}$

Specify fluid - assume isotropic, with no preferred axes. then  $\underline{q}$  must be parallel to  $\nabla T$ . Thus,

$$\underline{\kappa} = \kappa_{ij} = -k \delta_{ij} = -k \underline{I}$$

Minus sign convention so heat flux is positive in direction of decreasing T.

$$q_i = -k \delta_{ij} \frac{\partial T}{\partial x_j} = -k \nabla T$$

k is the thermal conductivity, can depend on local thermodynamic state (this is often neglected).

(2)  $\underline{\tau} = \tau_{ij}$ , the viscous stress tensor (Newtonian fluid)

$\underline{\tau}$  cannot depend on  $\underline{x}$  (translation of coords)

or on  $\underline{y}$  (Galilean invariance). Must depend on spatial derivatives of  $\underline{y}$ . Assume depends on 1st derivatives only. (like a Taylor series...). Assume a linear relation between  $\tau_{ij}$  and  $\frac{\partial u_i}{\partial x_j}$ .

For later assume that a rigid-body rotation does not produce viscous stresses. Thus,

decompose  $\frac{\partial u_i}{\partial x_j}$  into symmetric and anti-symmetric parts:

$$\frac{\partial u_i}{\partial x_j} = \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} - \frac{\partial u_j}{\partial x_i} \right)$$

but  $\epsilon_{ij} \equiv \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)$  rate of strain tensor

here wed 1-24.

$$\Omega_{ij} \equiv \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} - \frac{\partial u_j}{\partial x_i} \right) \Rightarrow \Omega_{ii} (\text{no sum}) = 0$$

$$\Omega_{ij} = -\Omega_{ji}$$

only 3 independent components  $\Rightarrow$  can associate a pseudovector with this tensor -

$$\rightarrow \begin{pmatrix} 0 & \Omega_{ij} \\ 0 & 0 \\ -\Omega_{ij} & 0 \end{pmatrix}$$

$$\Omega_{ij} = -\frac{1}{2} \epsilon_{ijk} \omega_k, \quad \underline{\omega} = \nabla \times \underline{u}$$

$\Omega_{ij}$  associated with rotation of fluid element. - rigid body rotations  $\Rightarrow$  no  $\tau_{ij}$   
 $\tau_{ij}$  depends only on  $\epsilon_{ij}$  - rate of strain tensor.

So  $\tau_{ij}$  is a linear function of  $\epsilon_{ij}$  - implies

$\epsilon_{ij}$  NOT permutation tensor now (only 2 indices)

$$\tau_{ij} = B_{ijkl} \epsilon_{kl}$$

81 components - and a fluid crystal can have many of them.

Each  $\tau_{ij}$  in general depends on every  $\epsilon_{kl}$ .

To simplify, assume fluid is isotropic. (water, air) (no preferred direction). Most general isotropic tensor

$$B_{ijkl} = \mu \delta_{ik} \delta_{jl} + \mu' \delta_{il} \delta_{jk} + \mu'' \delta_{ij} \delta_{kl}$$

( $\delta$  clearly isotropic) (can be shown, in tensor books)

But since  $\tau_{ij} = \tau_{ji}^*$ ,  $B_{ijkl} = B_{jilk} \Rightarrow \mu' = \mu$  (or get using symmetry of  $\epsilon_{ij}$ )

$$\tau_{ij} = [\mu(\delta_{ik} \delta_{jl} + \delta_{il} \delta_{jk}) + \mu'' \delta_{ij} \delta_{kl}] \epsilon_{kl}$$

$$\tau_{ij} = \mu(\epsilon_{ij} + \epsilon_{ji}) + \mu'' \delta_{ij} \epsilon_{kk}$$

$$\tau_{ij} = 2\mu \epsilon_{ij} + \mu'' \delta_{ij} \epsilon_{kk}$$

$\mu, \mu''$  arbitrary constants.

see Batchelor p143ff

\*otherwise  $\epsilon_{ij}$  rotational moments at a point.

$$\text{but } \epsilon_{kk} = \frac{1}{2} \left( \frac{\partial u_k}{\partial x_k} + \frac{\partial u_k}{\partial x_k} \right) = \frac{1}{2} \cdot 2 \nabla \cdot \underline{u}$$

$$\epsilon_{kk} = \nabla \cdot \underline{u}$$

$$\text{so } \tau_{ij} = 2\mu \epsilon_{ij} + \mu'' \delta_{ij} \nabla \cdot \underline{u}$$

$$= 2\mu \frac{1}{2} \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \mu'' \delta_{ij} \nabla \cdot \underline{u}$$

(index must be repeated in a single term for summ. conv.)

$$\tau_{ij} = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \delta_{ij} \mu'' \nabla \cdot \underline{u}$$

$$\text{Now trace } (\tau_{ij}) = \tau_{ii} = \mu \cdot 2 \cdot \frac{\partial u_i}{\partial x_i} + \delta_{ii} \mu'' \frac{\partial u_i}{\partial x_i}$$

$$= (2\mu + 3\mu'') \frac{\partial u_i}{\partial x_i} \equiv 3\lambda \frac{\partial u_i}{\partial x_i}$$

$$3\mu'' = 3\lambda - 2\mu$$

$$\text{so } \mu'' = \lambda - \frac{2}{3}\mu$$

$$\tau_{ij} = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + \delta_{ij} \left( \lambda - \frac{2}{3}\mu \right) \frac{\partial u_k}{\partial x_k}$$

See Rosenhead et al. in Proc. Roy Soc. A, 1954, pp. 1-6, re 2nd viscosity

see PFA5 1267, 1993

$\mu$  is ordinary viscosity.

$\lambda$  is the 'second viscosity' (often neglected) (important in shocks etc)

for second viscosity, see also notes from Erickson via Steinle, June 2014, plus AIAA paper 2004-0017 by Pan et al., etc. See also Meador et al., Phys. Fluids, v. 8, pp. 258-261, 1996, and Vincenti and Kruger, Intro. to Physical Gas Dynamics, 1965, pp. 407-412. Also the 23-page article by M.S. Cramer in Physics of Fluids, v. 24, 2012, article 066102. Remains controversial with many unresolved issues. Seem important for certain problems in ultrasound and perhaps second-mode instability waves in hypersonics.

③ Also note that if we give the equation of state in canonical form, we get the caloric equation at the same time - give

$$h(p, s) \quad (dh = Tds - vdp)$$

### J. Boundary Conditions

No-slip  $\rightarrow y_{fluid} \text{ at surface} = y_{boundary}$   
(both parallel and  $\perp$  components)

↙ molecule

on molecular scale, surface is not flat or hard. Bounces back randomly

really: particles (H<sub>2</sub>O) usually accepted and then reevaporated. Or a monolayer of adsorbed fluid.

If we neglect viscosity can't enforce both BC.  $\rightarrow$  because viscosity multiplies highest order derivative, if  $y=0$  lose one order in the equations.

at HWL  
19007  
p. 2  
notes.

FORM B  
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### k. Further Simplification of Momentum Equation

Insert the formula for the Newtonian shear into the momentum equation:

$$\frac{\partial(\rho u_i)}{\partial t} + \text{div}(\rho u_i u_j + p \delta_{ij} - \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) - \delta_{ij} (\lambda - 2/3 \mu) \frac{\partial u_k}{\partial x_k}) = f_i$$

or,

$$\frac{\partial(\rho u_i)}{\partial t} + \frac{\partial}{\partial x_j} \left[ \rho u_i u_j + p \delta_{ij} - \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) - \delta_{ij} (\lambda - 2/3 \mu) \frac{\partial u_k}{\partial x_k} \right] = f_i$$

$$\frac{\partial(\rho u_i)}{\partial t} + u_i \frac{\partial}{\partial x_j} (\rho u_j) + \rho u_j \frac{\partial}{\partial x_j} (u_i) + \frac{\partial p}{\partial x_i} - \frac{\partial}{\partial x_j} \left( \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right) - \frac{\partial}{\partial x_i} \left[ (\lambda - 2/3 \mu) \frac{\partial u_k}{\partial x_k} \right] = f_i$$

Abstract  
div (continuity)

If fluid is not incompressible, there is very little which can be done to simplify this equation.

$$\begin{aligned} \text{If } \mu = \text{const, } \frac{\partial}{\partial x_j} \left( \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right) &= \mu \frac{\partial^2 u_i}{\partial x_j \partial x_j} + \mu \frac{\partial^2 u_j}{\partial x_j \partial x_i} \\ &= \mu \nabla^2 u_i + \mu \nabla (\nabla \cdot u) \end{aligned}$$

(three terms total, in  $\nabla \cdot u$  or  $\nabla \cdot (\rho u)$ )

Continuity is

$$\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{y}) = 0$$

if  $\rho = \text{const}$  (incompressible, not stratified fluid) then  $\nabla \cdot \underline{y} = 0$

then, momentum equation becomes

$$\rho \frac{\partial u_i}{\partial t} + \rho u_j \frac{\partial}{\partial x_j} (u_i) = -\frac{\partial p}{\partial x_i} + \mu \frac{\partial^2 u_i}{\partial x_j \partial x_j} + f_i$$

or

$$\rho \frac{\partial \underline{y}}{\partial t} + \rho (\underline{y} \cdot \nabla) \underline{y} = -\nabla p + \mu \nabla^2 \underline{y} + \underline{f}$$

'Navier-Stokes' equation.

But this is true only if  $\rho = \text{const}$  (perhaps this can be relaxed a bit...)

9-794. L. Dissipation into heat, and the Entropy Equation

Energy Equation:

$$\frac{\partial (\rho (e + \frac{1}{2} u^2))}{\partial t} + \text{div} \left[ \rho \underline{y} (h + \frac{1}{2} u^2) + \underline{q} - \underline{\tau} \cdot \underline{y} \right] = \underline{f} \cdot \underline{y}$$

Momentum eqn:

$$\frac{\partial \rho \underline{y}}{\partial t} + \text{div} (\rho \underline{y} \underline{y} + p \underline{I} - \underline{\tau}) = \underline{f}$$



dot momentum eqn. into  $\underline{u}$  to get mechanical energy eqn:

$$\underline{u} \cdot \frac{\partial \rho \underline{u}}{\partial t} + \underline{u} \cdot \text{div} [e \underline{u} \underline{u} + p \underline{\underline{I}} - \underline{\tau}] = \underline{u} \cdot \underline{f}$$

want to subtract this from energy: Reformulate

$$\textcircled{1} \underline{u} \cdot \frac{\partial \rho \underline{u}}{\partial t} = u_j \frac{\partial \rho u_j}{\partial t} = \frac{\partial}{\partial t} (\rho u_j u_j) = \rho u_j \frac{\partial u_j}{\partial t} + u_j \frac{\partial}{\partial t} (\rho u_j)$$

$$\text{so } \underline{u} \cdot \frac{\partial \rho \underline{u}}{\partial t} = \frac{\partial}{\partial t} (\rho u^2) - \rho u_j \frac{\partial u_j}{\partial t} = \frac{\partial}{\partial t} (\rho u^2) - \rho \frac{\partial}{\partial t} \left( \frac{u^2}{2} \right)$$

$$\textcircled{2} \underline{u} \cdot \text{div} (\rho \underline{u} \underline{u}) = u_j \frac{\partial}{\partial x_j} (\rho u_i u_j) = u_j \frac{\partial}{\partial x_j} (\rho u_i) + u_j \rho u_i \frac{\partial u_j}{\partial x_j}$$

=

$$= u_i u_j \frac{\partial}{\partial x_j} (\rho u_i) + \rho u_j u_i \frac{\partial u_j}{\partial x_j}$$

~~energy eqns~~

$$\textcircled{3} \frac{\partial}{\partial t} (\rho e + \frac{1}{2} \rho u^2) = \frac{\partial}{\partial t} (\rho e) + \frac{1}{2} \frac{\partial}{\partial t} (\rho u^2)$$

$$\textcircled{4} \text{div} (\rho \underline{u} (h + \frac{1}{2} u^2)) = \frac{\partial}{\partial x_j} [\rho u_j (h + \frac{1}{2} u^2)] = \text{div} (\rho \underline{u} h) + \rho u_j \frac{\partial}{\partial x_j} (\frac{1}{2} u^2) + \frac{1}{2} u^2 \frac{\partial}{\partial x_j} (\rho u_j)$$

~~energy eqns~~

Subtract ① & ② from ③ & ④:

$$\frac{\partial}{\partial t}(\rho e) + \frac{1}{2} \frac{\partial}{\partial t}(\rho u^2) + \text{div}(\rho u h) + \cancel{(\rho u \nabla \cdot)} \left( \frac{1}{2} u^2 \right) + \frac{1}{2} u^2 \nabla \cdot (\rho u)$$

$$- \frac{\partial}{\partial t}(\rho u^2) + \rho \frac{\partial}{\partial t} \left( \frac{1}{2} u^2 \right) - u^2 \nabla \cdot (\rho u) - \cancel{(\rho u \nabla \cdot)} \left( \frac{u^2}{2} \right)$$

$$= \frac{\partial}{\partial t}(\rho e) + \text{div}(\rho u h) + \frac{1}{2} \frac{\partial}{\partial t}(\rho u^2) - \frac{1}{2} u^2 \nabla \cdot (\rho u) - \frac{\partial}{\partial t}(\rho u^2)$$

$$+ \rho \frac{\partial}{\partial t} \left( \frac{1}{2} u^2 \right)$$

$$= \frac{\partial}{\partial t}(\rho e) + \text{div}(\rho u h) - \frac{1}{2} \left[ \frac{\partial}{\partial t}(\rho u^2) + u^2 \nabla \cdot (\rho u) \right] + \rho \frac{\partial}{\partial t} \left( \frac{1}{2} u^2 \right)$$

$$\text{but } \frac{\partial}{\partial t}(\rho u^2) = u^2 \frac{\partial \rho}{\partial t} + \rho \frac{\partial u^2}{\partial t}$$

$$- \frac{1}{2} \frac{\partial}{\partial t}(\rho u^2) = - \frac{1}{2} u^2 \frac{\partial \rho}{\partial t} - \frac{1}{2} \rho \frac{\partial u^2}{\partial t}$$

$$= \frac{\partial}{\partial t}(\rho e) + \text{div}(\rho u h) - \frac{1}{2} \left[ u^2 \frac{\partial \rho}{\partial t} + u^2 \nabla \cdot (\rho u) \right] - \frac{1}{2} \rho \frac{\partial u^2}{\partial t} + \frac{1}{2} \rho \frac{\partial u^2}{\partial t}$$

$$\downarrow = u^2 (\text{continuity}) = 0$$

$$= \frac{\partial}{\partial t}(\rho e) + \text{div}(\rho u h)$$

So, energy eqn - mechanical energy eqn is:

$$\frac{\partial}{\partial t}(\rho e) + \text{div}(\rho \underline{y} h) + \text{div} \left[ \underline{q} - \underline{\tau} \cdot \underline{y} \right] - \underline{y} \cdot \text{div} \left[ \rho \underline{I} - \underline{\tau} \right] = \underline{f} \cdot \underline{y} - \underline{y} \cdot \underline{f} = 0$$

(8) of 16 Jan 06 notes

$$\frac{\partial}{\partial t}(\rho e) + \text{div}(\rho \underline{y} h + \underline{q} - \underline{\tau} \cdot \underline{y}) - \underline{y} \cdot \text{div}(\rho \underline{I} - \underline{\tau}) = 0$$

$$-\text{div}(\underline{\tau} \cdot \underline{y}) + \underline{y} \cdot \text{div} \underline{\tau} = -\frac{\partial}{\partial x_j} (\tau_{ij} u_j) + u_k \frac{\partial}{\partial x_j} (\tau_{ij})$$

$$= -\tau_{ij} \frac{\partial u_k}{\partial x_j} - u_k \frac{\partial \tau_{ij}}{\partial x_j} + u_k \frac{\partial \tau_{kj}}{\partial x_j} = -\tau_{ij} \frac{\partial u_k}{\partial x_j}$$

$$\underline{y} \cdot \text{div}(\rho \underline{I}) = u_k \frac{\partial}{\partial x_i} (\rho \delta_{ij}) = u_k \frac{\partial \rho}{\partial x_i} = \underline{y} \cdot \nabla \rho$$

$$\frac{\partial}{\partial t}(\rho e) + \text{div}(\rho \underline{y} h + \underline{q}) = \underline{y} \cdot \nabla \rho + \tau_{ij} \frac{\partial u_k}{\partial x_j}$$

$\phi = \tau_{ij} \frac{\partial u_k}{\partial x_j}$ , the dissipation of mechanical energy into heat through viscous shear.

$\underline{y} \cdot \nabla \rho$  = pressure work term.

$\text{div} \underline{q}$  = heat flux term.

here for: do 1-3  
1-26.

What is  $\phi$ ? (for a Newtonian fluid)

$$\phi = \tau_{ij} \frac{\partial u_i}{\partial x_j} = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \frac{\partial u_i}{\partial x_j} + \delta_{ij} (\lambda^{-2/3} \mu) \frac{\partial u_k}{\partial x_k} \frac{\partial u_i}{\partial x_j}$$

$$= \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \frac{\partial u_i}{\partial x_j} + (\lambda^{-2/3} \mu) \frac{\partial u_k}{\partial x_k} \frac{\partial u_i}{\partial x_i} \rightarrow \left( \frac{\partial u_k}{\partial x_k} \right)^2$$

Switch the dummy indices  $i, j$

$$\phi = \mu \left( \frac{\partial u_j}{\partial x_i} + \frac{\partial u_i}{\partial x_j} \right) \frac{\partial u_j}{\partial x_i} + (\lambda^{-2/3} \mu) \left( \frac{\partial u_k}{\partial x_k} \right)^2, \text{ add}$$

$$2\phi = \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) + 2(\lambda^{-2/3} \mu) \left( \frac{\partial u_k}{\partial x_k} \right)^2$$

Note:  
Batchelor  
p.153 seems  
to assume

$$\phi = \frac{1}{2} \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)^2 + (\lambda^{-2/3} \mu) \left( \frac{\partial u_k}{\partial x_k} \right)^2$$

Positive  
definite  
(unless  $\lambda^{-2/3} \mu$   
wired...?)

$\lambda = 2/3 \mu \rightarrow$  this is associated with his assumption

that  $\tau_{ii} = 0$  (sum of mean normal viscous stress = 0).

Not clear issue (hard to do expts, not important?)

Now, continue w/ derivation of entropy eqn.

$$\frac{\partial}{\partial t}(\rho e) + \text{div}(\rho \mathbf{y} h + \mathbf{q}) = \mathbf{y} \cdot \nabla p + \phi$$

$$\text{want } \frac{\partial}{\partial t}(\rho h) = \frac{\partial}{\partial t}(\rho(e + \frac{p}{\rho})) = \frac{\partial}{\partial t}(\rho e + p)$$

$$\frac{\partial}{\partial t}(\rho h) - \frac{\partial p}{\partial t} + \text{div}(\rho \mathbf{y} h) + \text{div} \mathbf{q} = \mathbf{y} \cdot \nabla p + \phi$$

$$\frac{\partial}{\partial t}(\rho h) + \text{div}(\rho \mathbf{y} h) - \left( \frac{\partial p}{\partial t} + \mathbf{y} \cdot \nabla p \right) = \phi - \text{div} \mathbf{q}$$

$$\rho \frac{\partial h}{\partial t} + h \frac{\partial \rho}{\partial t} + (\rho u_j \frac{\partial}{\partial x_j}) h + h \frac{\partial}{\partial x_j} (\rho u_j) - \left( \frac{\partial p}{\partial t} + \mathbf{y} \cdot \nabla p \right) = \phi - \text{div} \mathbf{q}$$

$$\rho \left[ \frac{\partial h}{\partial t} + (\mathbf{y} \cdot \nabla) h \right] + h \left[ \frac{\partial \rho}{\partial t} + \text{div}(\rho \mathbf{y}) \right] - \left( \frac{\partial p}{\partial t} + \mathbf{y} \cdot \nabla p \right) = \phi - \text{div} \mathbf{q}$$

(continuity)

Now derivative following the fluid is

$$\frac{d}{dt} = \frac{\partial}{\partial t} + \mathbf{y} \cdot \nabla \quad (\text{to be discussed later})$$

9-9-94

$$\text{so } \rho \frac{dh}{dt} - \frac{dp}{dt} = \phi - \text{div} \mathbf{q}$$

$$\text{but } T ds = dh - v dp = dh - \frac{1}{\rho} dp$$

$$\rho T ds = \rho dh - dp$$

so  $\rho T \frac{ds}{dt} = \rho \frac{dh}{dt} - \frac{dp}{dt}$  (tricky)

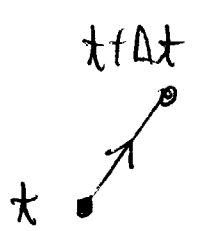
$$\rho T \frac{ds}{dt} = \phi - \text{div } q \quad (\text{cf. Batchelor } 304011)$$

$$\rho T \frac{ds}{dt} = \frac{1}{2} \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right)^2 + (\lambda - \frac{2}{3} \mu) \left( \frac{\partial u_k}{\partial x_k} \right)^2 - \text{div } q$$

neglected  
& viscous effects  
here  $dw = -pdv$ .  
diffusive  
effects here

Note Derivative following the fluid:

①



'particle' at  $(x, y, z)$  at time  $t$  moves to  $(x + \Delta x, y + \Delta y, z + \Delta z)$  at time  $t + \Delta t$

( $\phi$  any property carried w/ fluid)

$$\left. \frac{d\phi}{dt} \right|_{\text{following fluid}} = \frac{\phi(t + \Delta t) - \phi(t)}{\Delta t}, \Delta t \rightarrow 0$$

$$\left. \frac{d\phi}{dt} \right|_{f.f.} = \frac{\phi(t + \Delta t, x + \Delta x, y + \Delta y, z + \Delta z) - \phi(t)}{\Delta t}, \Delta t \rightarrow 0$$

but  $\phi(t + \Delta t, x + \Delta x) \approx \phi(t, x) + \frac{\partial \phi}{\partial t} \Delta t + \frac{\partial \phi}{\partial x} \Delta x + \frac{\partial \phi}{\partial y} \Delta y + \frac{\partial \phi}{\partial z} \Delta z + \dots$

so  $\left. \frac{d\phi}{dt} \right|_{f.f.} = \lim_{\Delta t \rightarrow 0} \frac{\partial \phi}{\partial t} + \frac{\partial \phi}{\partial x} \frac{\Delta x}{\Delta t} + \frac{\partial \phi}{\partial y} \frac{\Delta y}{\Delta t} + \frac{\partial \phi}{\partial z} \frac{\Delta z}{\Delta t} = \frac{\partial \phi}{\partial t} + (u \cdot \nabla) \phi$

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Note 2:  $s$  is the entropy per unit mass, so  $\rho s$  is the entropy per unit volume.  $\phi$  is the increase of entropy due to diffusion of momentum (randomness increases on molecular reduction of gradients).

Why  $\text{div } q$ ? Must be similar; if  $q = -k \nabla T$ ,  $k = \text{const}$ ,  
 $\text{div } q = -k \nabla^2 T$ ,  $-\text{div } q = k \nabla^2 T$ , so

$$\rho T \frac{ds}{dt} = \phi + k \nabla^2 T$$

& the similarity is apparent.  $k \nabla^2 T$  is reduction of gradients in  $T$  due to molecular diffusion.

Hard to see much more.

-9-3

### M. Vorticity

(see, e.g. pp ff in Rosenhead, by Lighthill)

"the news and bones of fluid motion" (Kücheman)

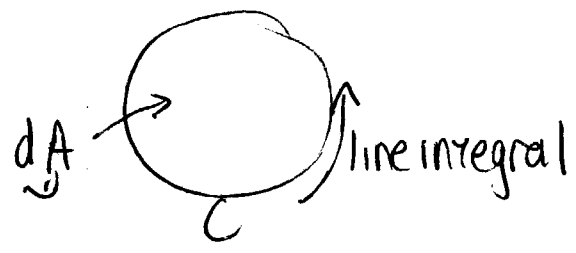
see  
Saffman,  
in HUL  
Festschrift  
'85

$$\underline{\omega} \equiv \nabla \times \underline{y} = \text{curl } \underline{y}$$

$$\int_S \underline{\omega} \cdot d\mathbf{A} = \int_C \underline{y} \cdot d\mathbf{l} \quad \text{by Green's theorem, (or Stokes),}$$

So vorticity associated with the rotation of a fluid element.

We have also seen this from the rotation tensor  $\Omega_{ij}$



Some theorems about vorticity:

① Helmholtz Law's: Vortex lines cannot end in the fluid:

since  $\nabla \cdot (\nabla \times \underline{u}) = 0$  for any vector field  $\underline{u}$ ,  
 $\nabla \cdot \underline{\omega} = 0$

A vortex line is a line everywhere tangent to the local vorticity vector (similar to a streamline). Since there is no divergence of  $\underline{\omega}$ , these lines cannot end in the fluid. In a viscous flow, they cannot end on the wall either; or else the no-slip condition would not be satisfied.

② Helmholtz Law #2: Consider a vortex tube, which is a tube with sides tangent to the vorticity (sides made up of vortex lines) so that  $\underline{\omega} \cdot \hat{n} = 0$  on the sides. What is the circulation around the ends?



Look at  $\int_A \underline{\omega} \cdot d\underline{A} = \int_A \underline{\omega} \cdot \hat{n} dS$   
 (A is a closed surface.)

$\int_A \underline{\omega} \cdot \hat{n} dS = \iiint_V \text{div } \underline{\omega} dV$  by divergence thm.

$\text{div } \underline{\omega} = \nabla \cdot (\nabla \times \underline{u}) = \frac{\partial}{\partial x_j} (\epsilon_{ijk} \frac{\partial}{\partial x_j} u_k) = 0$   
 $= \epsilon_{ijk} \frac{\partial^2 u_k}{\partial x_j \partial x_i}$



zero because  $j, k$  pairs cancel out, for each  $i$ .

thus,  $\int_A \underline{w} \cdot d\underline{A} = 0$ , for any flow.

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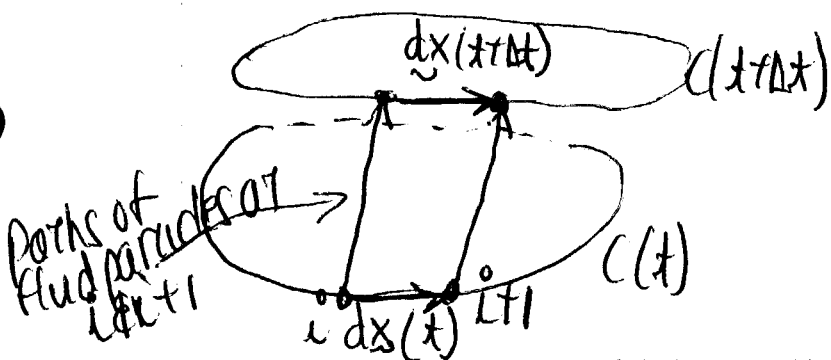
③ Kelvin's theorem: In an inviscid fluid in which  $\rho = \text{const}$  or  $\rho = \rho(p)$  only, and in which any body forces are conservative, vorticity moves with the fluid. (Action of viscosity pretty much restricted to diffusion).

Prove that  $\frac{d\Gamma}{dt} \Big|_{\text{moving w/ the fluid}} = 0$  with conditions above

$$\Gamma \equiv \int_C \underline{y} \cdot d\underline{x} = \iint_A \underline{w} \cdot d\underline{A}$$

$$\Gamma(t+\Delta t) - \Gamma(t) = \oint_{C(t+\Delta t)} \underline{y} \cdot d\underline{x} - \oint_{C(t)} \underline{y} \cdot d\underline{x}$$

where  $C$  marks the location of the same set of fluid particles, in both cases.



Note:  $\underline{y}_i \cdot \Delta \underline{x} + \underline{y}_i \cdot \underline{dx}(t+\Delta t)$   
 $= \underline{dx}(t) + \underline{y}_{i,t+\Delta t} \cdot \Delta t$   
(2 paths around 'square')

See AAE333  
Notes p.128ff

See  
Sommerfeld,  
Mechanics of  
Deformable  
Bodies,  
p.132ff.

$$\frac{d\Gamma}{dt} \Big|_{\text{moving w/ fluid}} = \lim_{\Delta t \rightarrow 0} \frac{\Gamma(t+\Delta t) - \Gamma(t)}{\Delta t}$$

$$= \frac{d}{dt} \Big|_{\text{m.w.f.}} \oint y \circ dx = \oint \frac{d}{dt} \Big|_{\text{m.w.f.}} (y \circ dx) \quad \left( \varepsilon \frac{d}{dx} = \frac{d}{dt} \varepsilon \right)$$

$$= \oint \left( \frac{dy}{dt} \Big|_{\text{m.w.f.}} \circ dx + y \circ \frac{d(dx)}{dt} \Big|_{\text{m.w.f.}} \right) = \oint \frac{dy}{dt} \Big|_{\text{m.w.f.}} \circ dx + \oint y \circ \frac{d(dx)}{dt} \Big|_{\text{m.w.f.}}$$

$$\text{but } \frac{d(dx)}{dt} \Big|_{\text{m.w.f.}} = \frac{dx(t+\Delta t) - dx(t)}{\Delta t}, \Delta t \rightarrow 0$$

$$= \frac{y_{t+\Delta t} \Delta t - y_t \Delta t}{\Delta t} \text{ (from } \square \text{)} = y_{t+\Delta t} - y_t$$

$$\text{so } \oint y \circ \frac{d(dx)}{dt} \Big|_{\text{m.w.f.}} \approx \sum_{i \text{ around loop}} y_i \circ (y_{t+\Delta t} - y_t)$$

$$\text{but } \sum_i y_i \circ (y_{t+\Delta t} - y_t) \approx \oint y \circ dy \quad \left( \begin{array}{l} \text{as } \Delta t \rightarrow \infty, \Delta y \rightarrow 0, \\ y_{t+\Delta t} - y_t = \Delta y \rightarrow dy \\ \varepsilon \rightarrow \int \end{array} \right)$$

$$\oint y \circ dy = \frac{1}{2} (y \circ y) \Big|_0^2$$

but for closed curve ① = ②,  $\oint y \circ dy = 0$ .

$$\left( \oint (u du + v dv + w dw) = \oint u du + \dots = \frac{1}{2} \oint d(u^2) + \dots = \frac{1}{2} y \circ y \right)$$

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here  
294.

so  $\frac{d\Gamma}{dt}|_{mwf} = \oint \frac{dy}{dt}|_{mwf} \cdot dx$

← go there 1-31

but  $\frac{dy}{dt}|_{mwf}$  is the 'material derivative'; use momentum eqns:

$$\frac{d}{dt}|_y = \frac{\partial y}{\partial t} + (y \cdot \nabla)y = -\frac{\nabla p}{\rho} + \frac{1}{\rho} \frac{\partial}{\partial x_j} \left[ \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right] + \frac{1}{\rho} \frac{\partial}{\partial x_k} (\lambda - \frac{2}{3}\mu) \frac{\partial u_k}{\partial x_k} + \frac{f}{\rho}$$

Terms in Integral:

a) if  $f$  is conservative,  $f = \nabla \psi$ ,  $\psi$  some potential.

For gravity,  $\frac{f}{\rho} = \nabla \phi$  (gravitational force proportional to mass)  
 ( $f$  is force per unit volume)

but  $\oint \nabla \phi \cdot dx = \phi|_1^2 = 0$  ( $\phi$  single-valued,  $\int_1^2 \nabla \phi \cdot dx = \phi|_1^2$ )

b)  $-\oint \frac{\nabla p}{\rho} \cdot dx$ ; if  $\rho = \text{const}$ , dispose of as above.

if  $\rho = \rho(p)$  only, then  $\frac{\nabla p}{\rho} = \nabla \left( \int \frac{dp}{\rho} \right)$ , and use same trick.

$$\left[ \frac{\partial p}{\partial x} \frac{1}{\rho} = \frac{\partial}{\partial x} \int \frac{dp}{\rho} = \frac{\partial}{\partial x} \left( \frac{dp}{\rho} \right) \right] \quad (?)$$

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When is  $\rho = \rho(p)$  only?

loose

$$T ds = de + p dv = de + p d\left(\frac{1}{\rho}\right)$$

$e$  depends on  $s$  and  $e$  &  $p$ , but  $e$  depends on  $p$  &  $T$ .

If  $s = \text{const}$  (inviscid & non-heat conducting)

then  $e$  depends on  $p$  &  $T$ . But we have the eqn of state to give  $T = T(e \& p)$  so  $e = e(p)$  only. (isentropic flow  $\Leftrightarrow \rho = \rho(p)$ )

and if inviscid, the other terms drop out too.

if not inviscid, 'errors' in Kelvin's theorem come from 'overturning' terms caused by density changes due to heating, and diffusion of vorticity caused by viscosity.

c) viscous terms: if  $\mu = \text{const}$ ,

$$\frac{1}{\rho} \frac{\partial}{\partial x_j} \left( \mu \left( \frac{\partial u_i}{\partial x_j} + \frac{\partial u_j}{\partial x_i} \right) \right) = \frac{\mu}{\rho} \left( \frac{\partial^2 u_i}{\partial x_j \partial x_j} + \frac{\partial^2 u_j}{\partial x_i \partial x_j} \right)$$

this term gives diffusion of vorticity, and stops the vortex tubes from moving only with the flow (think of diffusing line vortex).

$$\frac{1}{\rho} \delta_{ij} (\lambda + \frac{2}{3}\mu) \frac{\partial u_k}{\partial x_k}$$

?

So for conservative body forces, inviscid flow, and flow which is incompressible or isentropic, vorticity moves with the fluid (Kelvin's Theorem)

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④ Vorticity Equation

Start with momentum eqn:

$$\frac{\partial}{\partial t}(\rho \underline{y}) + \text{div}(\rho \underline{y} \underline{y} + p \underline{I} - \underline{\tau}) = \underline{f}$$

$$\rho \frac{\partial \underline{y}}{\partial t} + \underline{y} \frac{\partial \rho}{\partial t} + \text{div}(\rho \underline{y} \underline{y}) = \underline{f} + \text{div}(\underline{\tau} - p \underline{I})$$

$$\rho \frac{\partial u_i}{\partial t} + u_i \frac{\partial \rho}{\partial t} + \frac{\partial}{\partial x_j}(\rho u_i u_j) = \quad "$$

$$\rho \frac{\partial u_i}{\partial t} + \rho u_j \frac{\partial}{\partial x_j}(u_i) + \underbrace{u_i \frac{\partial \rho}{\partial t} + u_i \frac{\partial}{\partial x_j}(\rho u_j)}_{\rho \cdot (\text{continuity}) = 0} = \quad "$$

$$\rho \frac{\partial u_i}{\partial t} + \rho u_j \frac{\partial}{\partial x_j}(u_i) = \underline{f}_i + \text{div}(\underline{\tau}_i - p \underline{I}_i)$$

change form of left hand side:

$$\begin{aligned} \underline{\omega} \times \underline{y} &= \epsilon_{ilm} (\epsilon_{jsk} \frac{\partial}{\partial x_j} u_k) u_m \\ &= \epsilon_{ilm} \epsilon_{jsk} \left( \frac{\partial u_k}{\partial x_j} \right) u_m = (\delta_{mj} \delta_{sk} - \delta_{mk} \delta_{sj}) u_m \frac{\partial u_k}{\partial x_j} \end{aligned}$$

$$\omega \times y = u_j \frac{\partial u_k}{\partial x_j} - u_k \frac{\partial u_j}{\partial x_k} = u_j \frac{\partial u_k}{\partial x_j} - \frac{\partial}{\partial x_j} (u_k u_k \frac{1}{2})$$

$$\text{so } u_j \frac{\partial u_k}{\partial x_j} = \omega \times y + \nabla \left( \frac{u^2}{2} \right)$$

$$\rho \frac{\partial y}{\partial t} + \rho \omega \times y + \rho \nabla \left( \frac{u^2}{2} \right) = f + \text{div } \tau - \nabla p$$

alternative form of momentum eqn. - actually a correct vector form for l.h.s. - see Whitham, p. 120ff, in Rosentead (8p. 129)

can't make further progress without assuming incompressible. ( $\rho = \text{const}$ )

$$\text{then } \text{div } \tau = \mu \nabla^2 y$$

$$\frac{\partial y}{\partial t} + \omega \times y + \nabla \left( \frac{u^2}{2} + \frac{p}{\rho} \right) = \frac{1}{\rho} \left( f + \mu \nabla^2 y \right)$$

take the curl of this:

$$\frac{\partial \omega}{\partial t} + \nabla \times (\omega \times y) + \nabla \times \nabla \left( \frac{u^2}{2} + \frac{p}{\rho} \right) = \nabla \times \left( \frac{f}{\rho} \right) + \frac{\mu}{\rho} \nabla^2 \omega$$

0 if  $f$  conservative

$$\nabla \times (\omega \times y) = \epsilon_{ijk} \frac{\partial}{\partial x_j} (\epsilon_{klm} \omega_l y_m)$$

here for

$$= \epsilon_{kij} \epsilon_{klem} \frac{\partial}{\partial x_j} (\omega_e u_m)$$

$$= (\delta_{ie} \delta_{jm} - \delta_{im} \delta_{je}) \left( \frac{\partial}{\partial x_j} (\omega_e u_m) \right)$$

$$= \frac{\partial}{\partial x_j} (\omega_i u_j) - \frac{\partial}{\partial x_j} (\omega_j u_i)$$

$$= \omega_i \frac{\partial u_j}{\partial x_j} + u_j \frac{\partial \omega_i}{\partial x_j} - \omega_j \frac{\partial u_i}{\partial x_j} - u_i \frac{\partial \omega_j}{\partial x_j}$$

$$= \underline{\omega} \cdot \nabla \underline{u} + (\underline{u} \cdot \nabla) \underline{\omega} - (\underline{\omega} \cdot \nabla) \underline{u} - \underline{u} \cdot \nabla \underline{\omega}$$

but  $\nabla \cdot \underline{u} = 0$ , incompressible

$$\nabla \cdot \underline{\omega} = \nabla \cdot (\nabla \times \underline{u}) = \frac{\partial}{\partial x_j} (\epsilon_{jlm} \frac{\partial}{\partial x_l} u_m) = \epsilon_{jlm} \frac{\partial^2 u_m}{\partial x_j \partial x_l} = 0$$

$$\nabla \times (\underline{\omega} \times \underline{u}) = (\underline{u} \cdot \nabla) \underline{\omega} - (\underline{\omega} \cdot \nabla) \underline{u}$$

so

$$\frac{\partial \underline{\omega}}{\partial t} + (\underline{u} \cdot \nabla) \underline{\omega} = (\underline{\omega} \cdot \nabla) \underline{u} + \nu \nabla^2 \underline{\omega}$$

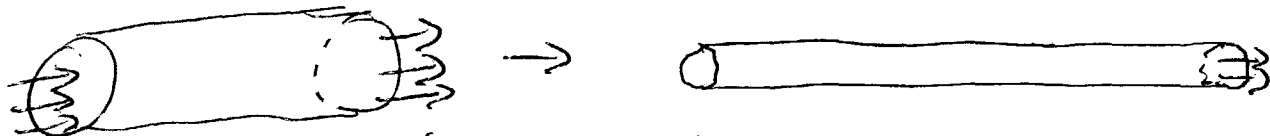
Vorticity Eqn.  
( $\rho = \text{const}$ , conservative  $f$ )

$$\frac{d \underline{\omega}}{dt} = (\underline{\omega} \cdot \nabla) \underline{u} + \nu \nabla^2 \underline{\omega}$$

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$\nu \nabla^2 \omega$  is the diffusion of vorticity term.

$(\omega \cdot \nabla) u$  is the vortex-stretching term.



if vortex tube lengthens, mass conservation implies that diameter must decrease, Helmholtz law  $\Rightarrow \omega$  increases. Like a spinning figure skater who pulls in her arms.

(see Batchelor or Whitham in Rosenhead for more details)