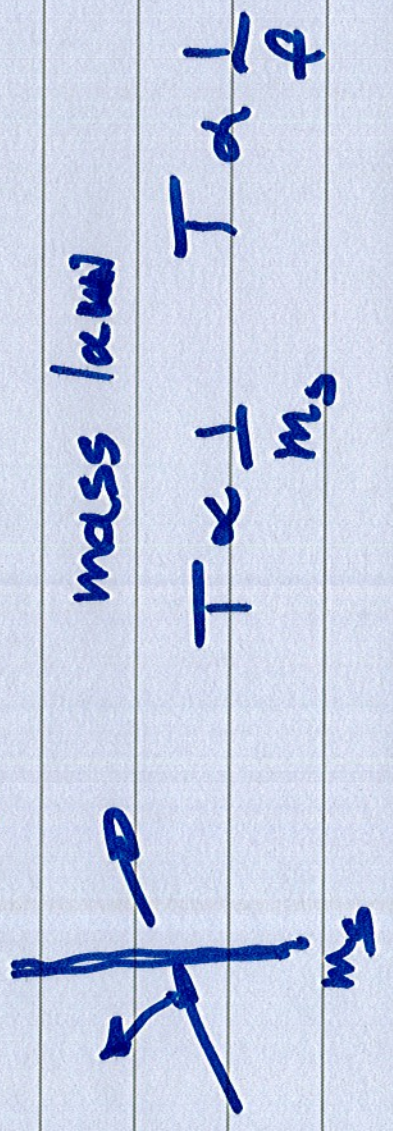
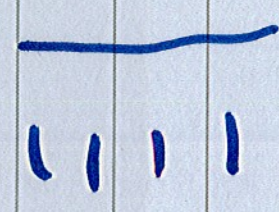
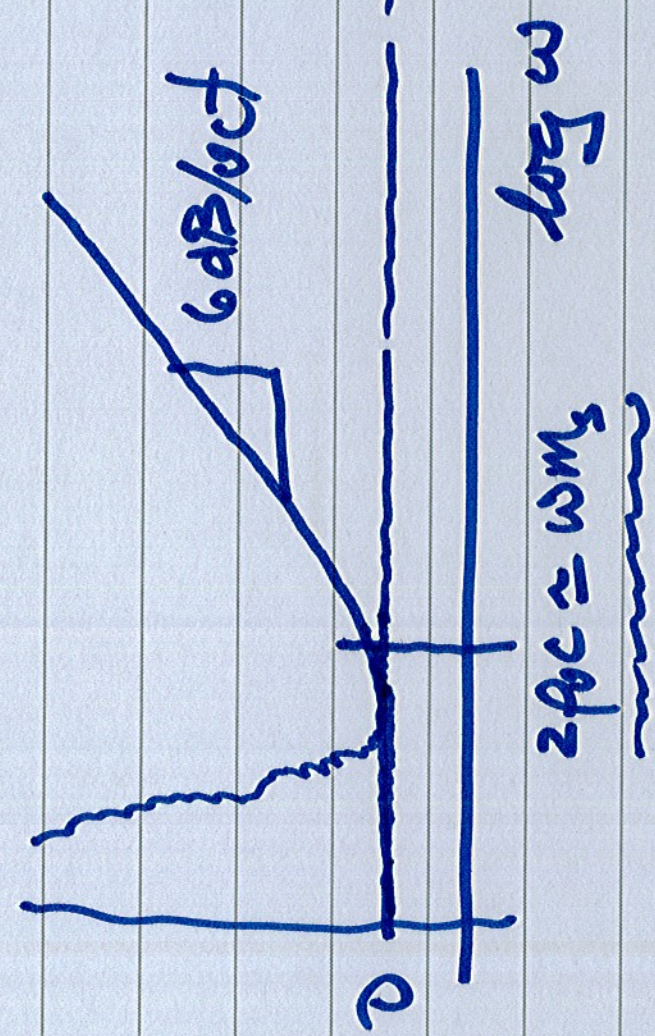
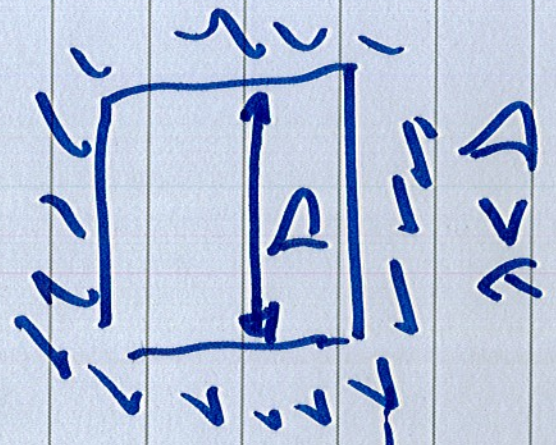


Same medium on both sides  $L = \frac{n\lambda}{2}$  Samar Dome

Thin lining barrier case

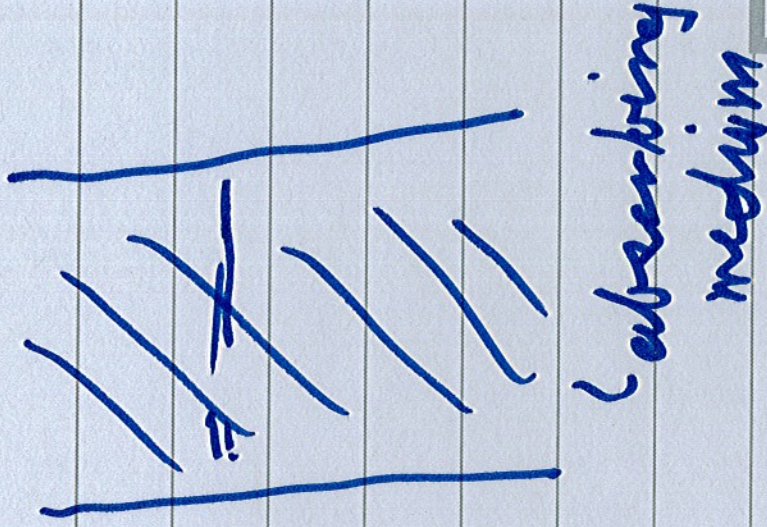




## Double Panel

$$TL \approx TL_1 + TL_2 + 6 \text{ dB}$$

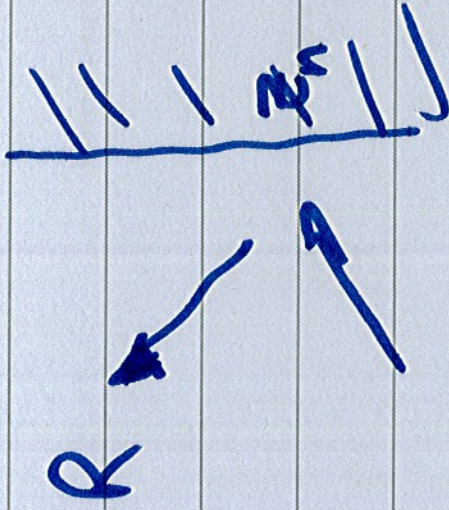
if there is adequate  
decoupling



- mass-air-mass



Surface with real impedance h.c.

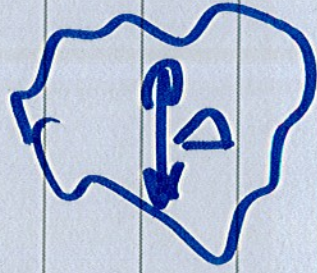


$$\alpha = 1 - |R|^2$$

# Section 5: Sound Generation & Radiation<sup>6</sup>

## Chapter 7

Compact Sources: small compared to a wavelength



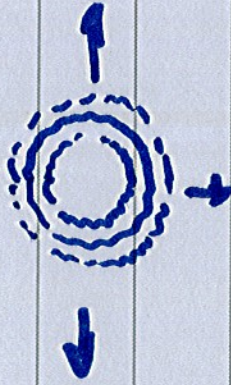
- loudspeakers at low frequency

- exhaust pipe

$$\underline{D \ll \lambda}$$

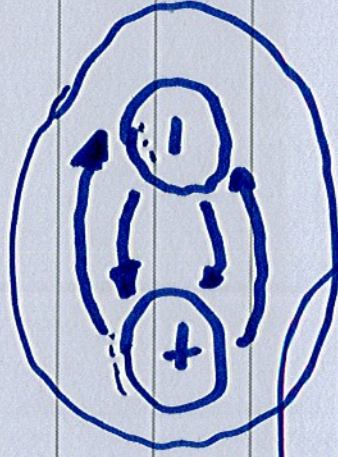
# "Simple" sources

- monopole - volume velocity source



- dipole

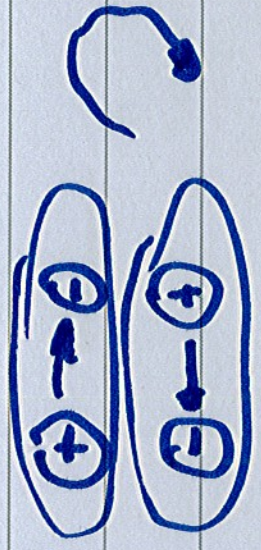
- Point force applied to the fluid



$\neq$  un baffled L/S

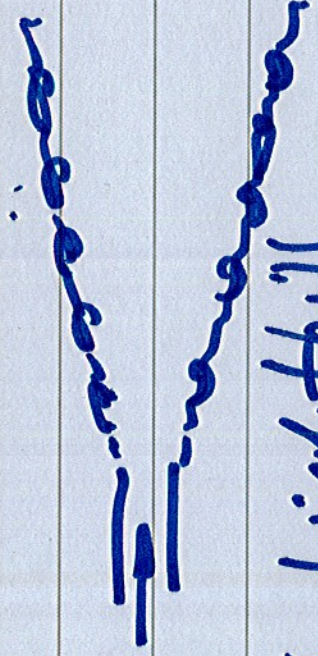
$\neq$  Axial fans

• Quadrupole



- an oscillatory moment applicable to fluid

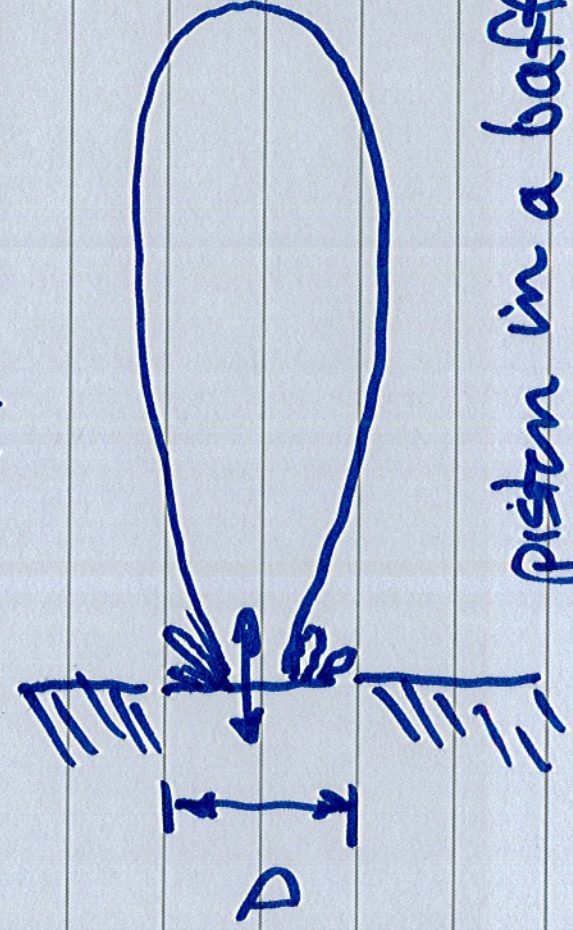
- turbulence jet



M. J. Lighthill  
1951

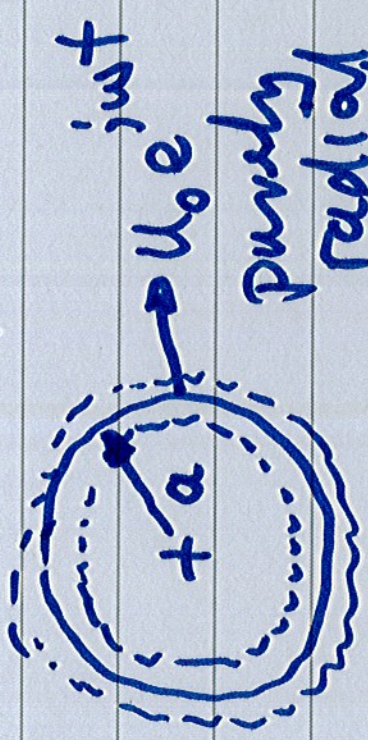


- Non-compact
- not small compared to  $\lambda$
- extended source



piston in a baffle

# S.2 Sound Radiation from a pulsating sphere



just

write:

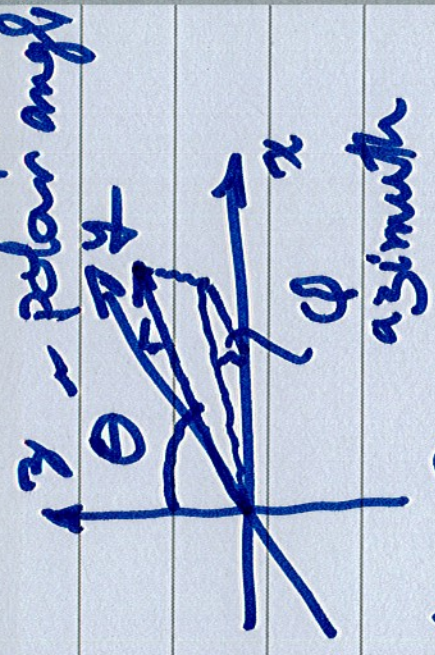
- wave eqn
- assume a solution
- apply a b.c at  $r=a$

purely radial velocity

spherically symmetric

Volume velocity source

no variation in  $\theta$  or  $\phi$   $\bar{p}(r)$



scalar Helmholtz Eqn harmonic

$$\nabla^2 \tilde{\psi} + k^2 \tilde{\psi} = 0$$

$$k = \frac{\omega}{c}$$

$$\psi(r, t) = \hat{\psi}(r) e^{j\omega t}$$

$$\nabla^2 \rightarrow \frac{\partial^2}{\partial r^2} + \frac{2}{r} \frac{\partial}{\partial r}$$

Due to the spherical symmetry

$$\frac{\partial}{\partial \phi} \rightarrow 0$$

$$\frac{d^2 \tilde{\psi}}{dr^2} + \frac{2}{r} \frac{d\tilde{\psi}}{dr} + k^2 \tilde{\psi} = 0$$

$$\frac{d^2(r\tilde{\psi})}{dr^2} + k^2(r\tilde{\psi}) = 0$$

$$\tilde{\psi} = \underbrace{\left(\frac{A}{r}\right) e^{-ikr}}_{\text{outward}} + \underbrace{\left(\frac{B}{r}\right) e^{+ikr}}_{\text{inward-going}}$$

neglect

Free-space - no reflecting surfaces

outward going wave only.

## 5.2.2 Boundary Conditions

set radial particle velocity  
at  $r = a$  to  $U_0 e^{i\omega t}$

$$\underbrace{\tilde{u}_r(a)}_{\text{fluid velocity}} = U_0 \underbrace{e^{i\omega t}}_{\text{surface velocity}}$$

$$\tilde{u}_r = -\frac{1}{j\omega b} \nabla^2 \tilde{p}$$

- no pressure gradients  
in  $\theta$  or  $\phi$  dir