

ASPHERICAL FORMATIONS NEAR THE LIBRATION POINTS IN THE SUN-EARTH/MOON SYSTEM B.G. Marchand and K.C. Howell Purdue University



Output Feedback Linearization (OFL) in the Ephemeris Model

- Formation Keeping + Deployment
 - Chief S/C Evolves Along Lissajous Trajectory near L_i
 - Inertial Formation Geometry
 - Spherical Configurations
 - Deputy Constrained to Orbit Chief S/C
 - Fixed Radial Distance
 - Fixed Radial Distance + Rotation Rate
 - Aspherical Configurations \rightarrow Inertially Fixed Orientation
 - Deputy Constrained to Evolve Along Aspherical Surface
 - Surface may be Offset from Chief S/C

OFL Controlled Response of Deputy S/C Radial Distance Tracking





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Impact of <u>Nominal</u> Radial Separation on OFL Controlled Response of Deputy S/C



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OFL Controlled Response of Deputy S/C Radial Distance + Rotation Rate Tracking



Impact Commanded Rotation Rate on Cost



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Application to Aspherical Formations

Parameterization of Parabolic Formation



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Controller Development

Desired Response for u, q, and \dot{v} :

$$g_{u}\left(u_{p},\dot{u}_{p}\right) = \ddot{u}_{p}^{*} - 2\omega_{n}\left(\dot{u}_{p}-\dot{u}_{p}^{*}\right) - \omega_{n}^{2}\left(u_{p}-u_{p}^{*}\right) \\ g_{q}\left(u_{p},\dot{u}_{p}\right) = \ddot{q}^{*} - 2\omega_{n}\left(\dot{q}-\dot{q}^{*}\right) - \omega_{n}^{2}\left(q-q^{*}\right) \\ g_{v}\left(\dot{v}\right) = \ddot{v}^{*} - k\omega_{n}\left(\dot{v}-\dot{v}^{*}\right) \\ \end{bmatrix} \delta\dot{\theta} \rightarrow \text{exponential decay}$$

Solve for Control Law:

$$\begin{bmatrix} \frac{2h}{a^2}\tilde{x} & \frac{2h}{a^2}\tilde{y} & 0\\ -\frac{2h}{a^2}\tilde{x} & -\frac{2h}{a^2}\tilde{y} & 1\\ \frac{\tilde{x}}{\left(\tilde{x}^2+\tilde{y}^2\right)} & -\frac{\tilde{y}}{\left(\tilde{x}^2+\tilde{y}^2\right)} & 0 \end{bmatrix} \begin{bmatrix} \tilde{u}_x\\ \tilde{u}_y\\ \tilde{u}_z \end{bmatrix} = \begin{bmatrix} g_u(u,\dot{u}) - \frac{2h}{a^2}\left(\dot{\tilde{x}}^2+\dot{\tilde{y}}^2+\tilde{x}\Delta\tilde{f}_x+\tilde{y}\Delta\tilde{f}_y\right) \\ g_q(q,\dot{q}) + \frac{2h}{a^2}\left(\dot{\tilde{x}}^2+\dot{\tilde{y}}^2+\tilde{x}\Delta\tilde{f}_x+\tilde{y}\Delta\tilde{f}_y\right) - \Delta\tilde{f}_z \\ g_v(\dot{v}) + 2\frac{\left(\tilde{x}\dot{\tilde{x}}+\tilde{y}\dot{\tilde{y}}\right)\left(\tilde{x}\dot{\tilde{y}}-\tilde{y}\dot{\tilde{x}}\right)}{\left(\tilde{x}^2+\tilde{y}^2\right)^2} + \frac{\left(\tilde{y}\Delta\tilde{f}_x-\tilde{x}\Delta\tilde{f}_y\right)}{\left(\tilde{x}^2+\tilde{y}^2\right)} \end{bmatrix}$$

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OFL Controlled Parabolic Formation



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OFL Thrust Profile



Original Maxim Design



http://maxim.gsfc.nasa.gov/documents/SPIE-2002/spie2002.ppt

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New Maxim Pathfinder



http://maxim.gsfc.nasa.gov/documents/SPIE-2002/spie2002.ppt



Maxim Configuration Example



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Building Non-Natural Formations Using Naturally Existing Solutions & Impulsive Control

Natural Formations: Quasi-Periodic Relative Orbits \rightarrow 2-D Torus





Natural Formations: Nearly Periodic + Slowly Expanding Orbits



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Evolution of Nearly Vertical Orbit Over 100 Orbital Periods



Enforcing Periodicity in the Ephemeris Model



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Geometry of Natural Solutions in the Ephemeris Model

Rotating Frame Perspective:

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Inertial Frame Perspective:





Concluding Remarks

OFL Control

- Successful for Formation Keeping & Reconfiguration
 - Spherical + Parabolic Formations
 - ↑ Rotation Rate = ↑ Thrust Level
 - w/o Rotation Rate \rightarrow Thrust ~ O(nN)
 - w/ Rotation Rate \rightarrow Thrust ~ O(mN)
- Natural to Non-Natural Formations
 - Differential Corrector \rightarrow 1 small maneuver/year
 - Can work well in the rotating frame
 - Depends on Impact of SRP
 - Difficult in the inertial frame due to Geometry of Initial Guess