IMPLICIT PRIORS FOR MODEL-BASED INVERSION

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ABSTRACT

While Markov random field (MRF) models have been widely used in the solution of inverse problems, a major disadvantage of these models is the difficulty of parameter estimation. At its root, this parameter estimation problem stems from the inability to explicitly express the joint distribution of an MRF in terms of the conditional distributions of elements given their neighbors. The objective of this paper is to provide a general approach to solving maximum a posteriori (MAP) inverse problems through the implicit specification of a MRF prior. In this method, the MRF prior is implemented through a series of quadratic surrogate function approximations to the MRF's log prior distribution. The advantage of this approach is that these surrogate functions can be explicitly computed from the conditional probabilities of the MRF, while the explicit Gibbs distribution can not. Therefore, the Gibbs distribution remains only implicitly defined. In practice, this approach allows for more accurate modeling of data through the direct estimation of the MRF's conditional probabilities. We illustrate the application of our method with a simple experiments of image denoising and show that it produces superior results to some widely used MRF prior models.

Index Terms— Markov random fields, Inverse problems, Maximum a posteriori estimation.

1. INTRODUCTION

Model-based inversion methods, first introduced decades ago for the solution of ill-posed inverse problems [1] have continued to gain importance as their value in the solution of difficult and widely used inverse problems grows [2, 3, 4]. A classical approach to model-based inversion is the computation of the maximum a posteriori (MAP) estimate which is given by

$$\hat{x} = \arg\max_{x \ge 0} \left\{ \log p(y|x) + \log p(x) \right\} ,$$

where p(y|x) is the forward model of the data vector y given the unknown vector x, and p(x) is the prior model for x. In many important applications, x is an image or 3D volume and p(x) is a prior model for the image.

Model-based inversion approaches have been enormously valuable in applications such as 3D reconstruction from X-ray computed tomography (CT) data [5]. A great advantage of model-based inversion methods is that they allow for the explicit incorporation of a forward model. However, a challenge is that they require the adoption of an explicit, tractable, and accurate prior model p(x). Perhaps

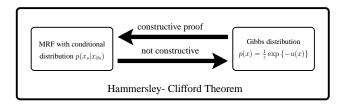


Fig. 1. The Hammersley-Clifford theorem states that X is an MRF with a strictly positive density if and only if it has a Gibbs distribution. However, the value of this very important theorem is limited by the fact that there is generally no tractable method to construct the Gibbs distribution from the conditional distributions.

the most common choice of prior model has been the Markov random field (MRF) because it limits the dependencies in x so that

$$p_{\theta}(x_s|x_r \ r \neq s) = p_{\theta}(x_s|x_{\partial s})$$
,

where x_s is an element of x, ∂s is the index set for the neighbors of s, and θ is a parameter vector which can be used to fit the prior distribution for specific application at hand.

Unfortunately, the conditional distribution $p_{\theta}(x_s|x_{\partial s})$ does not provide an explicit form for the prior distribution $p_{\theta}(x)$ of the associated MRF. The partial solution to this dilemma comes from the celebrated Hammersley-Clifford Theorem [6] which states that X is an MRF² if and only if its distribution can be expressed as a Gibbs distribution

$$p(x) = \frac{1}{z_{\theta}} \exp\{-u_{\theta}(x)\}\$$

where $u_{\theta}(x)$ is the Gibbs energy function which is formed by a sum of potential functions over neighborhood cliques, and z_{θ} is the so-called partition function. However, Figure 1 illustrates the limitation of the Hammersley-Clifford theorem. While the proof that a Gibbs distribution is an MRF is constructive, the proof that the MRF must have a Gibbs distribution with potential functions limited to cliques is not. Therefore, the Hammersley-Clifford Theorem does not provide a general trackable method for constructing the Gibbs distribution from the known conditional distribution, $p_{\theta}(x_s|x_{\partial r})$. It simply says that such a mapping exists.

The objective of this paper is to introduce a novel method for the construction of prior models which is based on direct estimation of the conditional distribution, $p_{\theta}(x_s|x_{\partial s})$, of a MRF model. The advantage of this approach is that it offers the opportunity to construct

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¹This is in contrast to a Markov chain in which the conditional probabilities can be simply multiplied to form the prior distribution.

 $^{^2}$ Technically, it must also have a strictly positive density, but in practice some values of X can occur with extremely low probability.

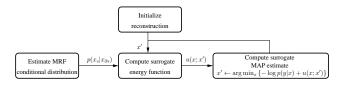


Fig. 2. Flow diagram illustrating the operations in MAP estimation with an implicit prior.

much more informative and accurate prior models through the use of a wide array of modern methods for the accurate estimation of conditional densities from training data. This is in contrast to traditional MRF approaches which typically assume a Gibbs distribution with simple potential functions controlled by a small number of parameters.

In this paper, we introduce a novel approach to computing the MAP estimate which does not require the Gibbs distribution to be explicitly computed. Instead, we compute local successive approximations to the energy function $u_{\theta}(x)$ using a surrogate energy function u(x; x') where x' is the point of approximation. As it turns out, this surrogate energy function can be explicitly computed from the MRFs conditional probabilities, whereas the underlying Gibbs distribution remains implicit. The surrogate energy function works in the same way as widely used majorization methods for optimization [7, 8, 9, 10, 11]. The key novelty to our method is in how we compute the surrogate energy function's form from the known (e.g. typically estimated from training data) conditional probabilities, $p_{\theta}(x_s|x_{\partial s})$. Using this method, we can then compute the MAP inversion with a nested iteration: Each "outer loop" updates the point of approximation x', and each "inner loop" maximizes a surrogate MAP optimization problem with a quadratic prior term.

2. BAYESIAN INVERSION USING IMPLICIT PRIOR

Our approach to Bayesian inversion with an implicit prior is illustrated in Figure 2. First, a the conditional distribution, $p_s(x_s|x_{\partial s})$, is either estimated from training data or simply selected based on knowledge of the application. The challenge is then to use this conditional density function as the basis for a prior model in MAP inversion. To do this, we will use the known conditional density to compute a surrogate energy function, u(x,x'), with the properties that

$$u(x') = u(x'; x') \tag{1}$$

$$u(x) < u(x; x'), \tag{2}$$

where u(x) is the energy function of the MRF's unknown Gibbs distribution. Once we obtain this surrogate energy function, we may use it to iteratively minimize the MAP cost function with the following procedure.

MAP estimation with implicit prior

- 1. Initialize x'
- 2. Repeat until x' has converged {
 - (a) Update surrogate energy function u(x; x')
 - (b) $x' \leftarrow \arg\min_{x} \left\{ -\log p(y|x) + u(x;x') \right\}$

}

Based on the theory of majorization, it is well known that the two properties of equations (1) and (2) insure monotone convergence of the MAP cost function with each iteration of the algorithm.

However, the problem remains of how to determine the surrogate energy function, u(x;x'), from the known conditional densities, $p_s(x_s|x_{\partial s})$. To do this, we first adopt a surrogate energy function with the form

$$u(x;x') = \frac{1}{2}(x-x')^t B(x-x') + d^t(x-x') + c, \quad (3)$$

where B is a symmetric matrix, d a column vector, and c a scalar, all of which are assumed to be functions of x'. We note that the quadratic form of the function makes solution of the surrogate MAP optimization problem straightforward using a wide range of standard optimization methods. Without loss of generality, we can simply assume that c=0 because the value of c does not affect the result of optimization.

The values of the parameter vector d can be easily computed from the gradient of the conditional densities as

$$d_s = -\left. \frac{\partial}{\partial x_s} \log p_s(x_s | x_{\partial s}) \right|_{x=x'} . \tag{4}$$

It only remains to choose the symmetric matrix B sufficiently large so that u(x;x') upper bounds the true energy function. However, if B is chosen to be too large, then convergence of the algorithm will be slow; so it is best to select a B which represents as tight an upper bound as possible. Therefore, our approach is to first find strong necessary conditions that B must satisfy, and then present a method to compute a matrix B that satisfies these conditions for our specific choice of the conditional distribution. Once this is done, we can then scale the magnitude of B or its diagonal as is necessary to insure an upper bound.

The following three conditions must hold for any matrix B which satisfies the equations of (1) and (2). (See proofs in [12].)

Condition 1: The symmetric matrix B must be positive definite. Condition 2: It must be the case that $B \geq H$, i.e. B-H must be a positive semi-definite matrix, where H is the Hessian of the energy function u(x) at x=x'. Moreover, the elements of H are given by

$$H_{s,r} = -\frac{\partial^2}{\partial x_s \partial x_r} \log p_s(x_s | x_{\partial s}) \bigg|_{x = x'} . \tag{5}$$

Condition 3: It must be the case that $B_{s,s} \geq D_{s,s}$ where D is a diagonal matrix with entries

$$D_{s,s} = 2 \sup_{x_s \neq x_s'} \left\{ \frac{-\log p_s(x_s | x_{\partial s}') + \log p_s(x_s' | x_{\partial s}') - \Delta_s d_s}{\Delta_s^2} \right\}$$
(6)

where $\Delta_s = x_s - x_s'$ and d_s is from equation (4). Furthermore, it is the case that $D_{s,s} \geq H_{s,s}$.

For our particular example, we will use a homogeneous conditional distribution based on the a Gaussian mixture distribution (See details in [12]). In this case, the conditional distribution has the form

$$p(x_s|x_{\partial s}) = \sum_k \frac{\gamma_k}{\sqrt{2\pi}\sigma_k} \exp\left\{-\frac{1}{2\sigma_k^2}(x_s - A_k x_{\partial s} - \beta_k)^2\right\},\,$$

where A_k is a row vector, β_k and σ_k are constants, and $\gamma_k = p(k|x_{\partial s})$. From this we can compute,

$$d_s = \sum_k \frac{1}{\sigma_k^2} (x_s' - A_k x_{\partial s}' - \beta_k) p(k|x_s', x_{\partial s}')$$
 (7)

$$H_{s,r} \stackrel{\simeq}{=} \tilde{H}_{s,r} = \sum_{k} \frac{1}{\sigma_k^2} (\delta_{s=r} - \delta_{s \neq r} A_{k,r}) p(k|x_s', x_{\partial s}') , (8)$$

where the Hession is efficiently computed using the approximation that $\frac{\partial}{\partial x_r} p(k|x_s, x_{\partial s})$ is small. We also note that it may be the case that the computed value of \tilde{H} is not symmetric, so we impose symmetry by computing $\tilde{H} \leftarrow (\tilde{H} + \tilde{H}^t)/2.^3$

Using the 1D case of the lemma proved in [12], it can be shown that $\tilde{H}_{s,s} \geq D_{s,s}$; so Condition 3 is met. It is important that B be positive definite (i.e. that Condition 1 holds) in order to insure that the inner loop is convergent. If this is not the case, then we can enforce Condition 1 by selecting B so that

$$B = \tilde{H} + \alpha \operatorname{diag}\{\tilde{H}\}, \tag{9}$$

where $\operatorname{diag}\{\tilde{H}\}$ is the positive-definite matrix formed by the diagonal of \tilde{H} , and $\alpha \geq 0$ is a positive constant. Notice that from the form of (9) and with the approximation of (8), then Condition 2 is met.

3. EXPERIMENTAL RESULTS





(a) Ground truth image x

(b) Noisy image y

Fig. 3. (a) Ground truth image, (b) image with additive white noise standard deviation $\sigma_w = 20$

In order to compare the implicit prior method against alternative prior models, we performed simulations for the very simple inverse problem of removing additive white Gaussian noise from an image [13]. More specifically, we generated a noisy observed image, y, from the "ground truth" image x by

$$y = x + w \;,$$

where w is i.i.d. Gaussian noise with distribution $N(0,\sigma_w^2)$ with $\sigma_w=20$. Figure 3 shows the ground truth image x, noisy image y. The 25 grayscale training images were taken from a set of natural scene photos, and the ground truth image used in testing was not contained in the set of training images. The images were captured by a Nikon D90 camera, the RGB values were converted to the luma values, and then the images were filtered and subsampled down to approximately 363×288 resolution, so as to be most suitable for illustrating results in this publication. Parameters of the MRF conditional probability density, $p(x_s|x_{\partial s})$, were estimated as described in [14] using M=32 and p=0.50.

In addition to using the implicit prior, we also ran comparisons with a range of different parameter values for the generalized Gaussian MRF (GGMRF) [15], and with the more general qGGMRF [5] which represents the current state-of-the-art in MRF priors for inverse problems such as tomography. The energy function of the form

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$$u(x) = \frac{1}{P\sigma^P} \sum_{\{i,j\}} b_{i-j} \rho(\Delta)$$
 (10)

$$\rho(\Delta) = \frac{|\Delta|^P}{1 + |\frac{\Delta}{C}|^{P-Q}} \tag{11}$$

where $\Delta=x_i-x_j$, and the parameter constraints are $1\leq Q\leq P\leq 2$ and C is a positive threshold. The sum is over all pairs $\{i,j\}$ such that i and j are 8-point neighbors and the coefficients b_i sum to 1, and the ratio of $b_i/b_j=\sqrt{2}$ when i is a 4-point neighbor and j is an 8-point neighbor. We used P=2 and the two values of Q=1.0 and Q=1.2 to illustrate a typical non-Gaussian MRF prior. If we set P=Q, then this is the form of the generalized Gaussian MRF (GGMRF). In all cases, we adjusted the scale parameter, σ , to minimize the root mean squared error (RMSE) between the reconstruction and the ground-truth image.

Figure 4 shows the root mean squared error (RMSE) of the implicit prior method as a function of the number of iterations. The blue line shows the RMSE between the restored image and the ground-truth image; and the black lines shows the RMSE between the restored image and the converged result of the algorithm. The plot indicates that the MAP estimate converges after about 20 iterations. In practice, smaller values of α tend to result in faster convergence; however, if α is chosen to be too small then the convergence may not be robust. In this case, $\alpha=0.45$ was larger than necessary, but we found this value consistently produced robust convergence in a wide array of examples.

Figure 5 attempts to graphically illustrate the values in the matrix $B_{s,r}$ after 20 iterations. For a pixel s, the color was set to green $=255*B_{s,s+(0,-1)}/B_{s,s}$ and red = blue $=255*B_{s,s+(-1,0)}/B_{s,s}$, where r=s+(0,-1) is the pixel immediately to the left of s, and r=s+(-1,0) is the pixel immediately above s. This image shows how the local weights in the surrogate energy function adapt to the local edge structure in the image.

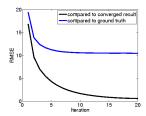




Fig. 4. Convergence of RMSE

Fig. 5. The matrix $B_{s,r}$ entry.

Figure 6 shows the comparison of MAP image reconstructions using the implicit prior and the qGGMRF prior with (P=2.0,Q=1.2) in Figure 6(c), and (P=2.0,Q=1.0) in Figure 6(d). In each case, the threshold C and the regularization σ were chosen to achieve the minimum RMSE. The implicit prior result is slightly sharper with slightly better detail than both the qGGMRF cases. This conclusion is supported by the objective measures of RMSE presented in Figure 7 where the implicit prior technique has the smallest value among the three techniques. The RMSE of the qGGMRF prior technique is plotted as a function of the threshold C while the RMSE of GGMRF prior technique is plotted as a function of P values. Notice that the result of C=0 for the qGGMRF with P and Q is equivalent to the result of P=Q for GGMRF because in both these cases the prior corresponds to the prior term $\rho(\Delta)=|\Delta|^P$,

³Theoretically, the matrix H should be symmetric. However, in practice the estimated conditional distribution may not correspond to a consistent MRF model, so it may be the case that for the Hession computed using equation (5), $H_{s,r} \neq H_{r,s}$.



(a) Ground truth



(c) qGGMRF (P = 2, Q = 1.2)



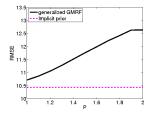
(b) Implicit prior

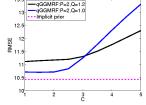


(d) qGGMRF (P = 2, Q = 1.0)

Fig. 6. The image restoration comparison between the implicit prior technique and qGGMRF prior technique ($P=2,\,Q=1.2,\,C=0$) and ($P=2,\,Q=1.0,\,C=1.5$). In each case, the threshold C and regularization parameter were chosen to achieve the minimum RMSE.

within a multiplicative constant. Interestingly, the total-variation prior does not produce the minimum RMSE reconstruction, but it does achieve a value near the minimum .





(a) Comparison with GGMRF

(b) Comparison with qGGMRF

Fig. 7. These plots compare the RMSE of the implicit prior with the RMSE of the GGMRF and qGGMRF priors, respectively. Each plot varies a parameter of the model. In each case, the regularization parameter was chosen to achieve the minimum RMSE.

4. CONCLUSIONS

We introduce a new method of MAP inversion which allows for the use of a Gibbs distribution which is only implicitly specified through the conditional probabilities of an MRF. The advantage of this approach is that it allows for the use of a much wider range of MRF models, which can in turn allow for more accurate modeling of data. The key to our approach is a method for explicitly computing local approximation to the energy function of the Gibbs distribution. These local approximations serve as a series of surrogate energy functions in the computation of the MAP inversion.

We provide a simple example of image denoising, but the method is generally applicable to any continuously valued MRF prior model, and could be combined with more sophisticated methods for estimation of the MRF's conditional distribution. Furthermore, the method has the potential for continuous improvement as methods for estimation of the MRF's conditional densities improve.

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